

RHIC RESULTS:

The Search for High Density Matter

What are we trying to understand?

What have we already learned?

What do we expect to learn?

What do we hope to learn?

What Are We Trying to Understand?

What is the behavior of matter at asymptotic energy density?

Quark Gluon Plasma

$$E/V > 1 \text{ GeV/Fm}^3$$

Early universe, neutron stars

What is the matter important for high energy hadrons?

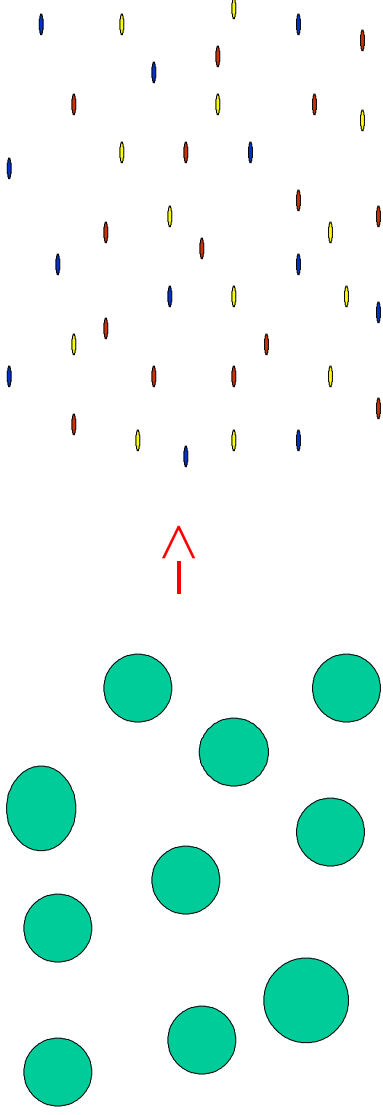
Color Glass Condensate

$$E/A > 1 \text{ GeV/Fm}^2$$

Universal high energy limit of strong interactions

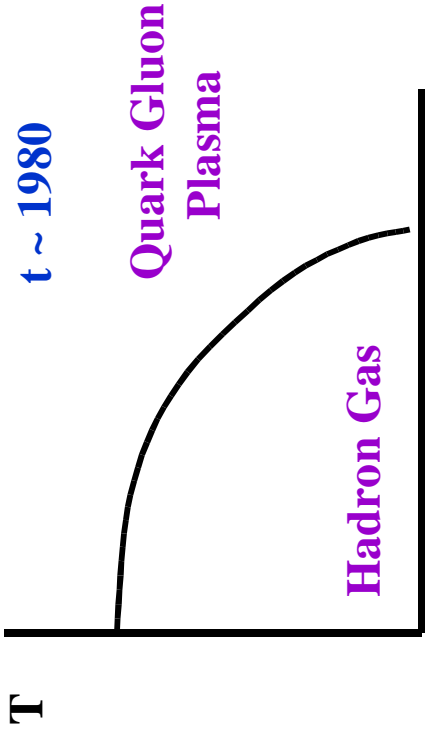
The Quark Gluon Plasma

Nucleons, mesons \rightarrow quarks, gluons



Low Energy Density \rightarrow High Energy Density

The Evolving Phase Diagram

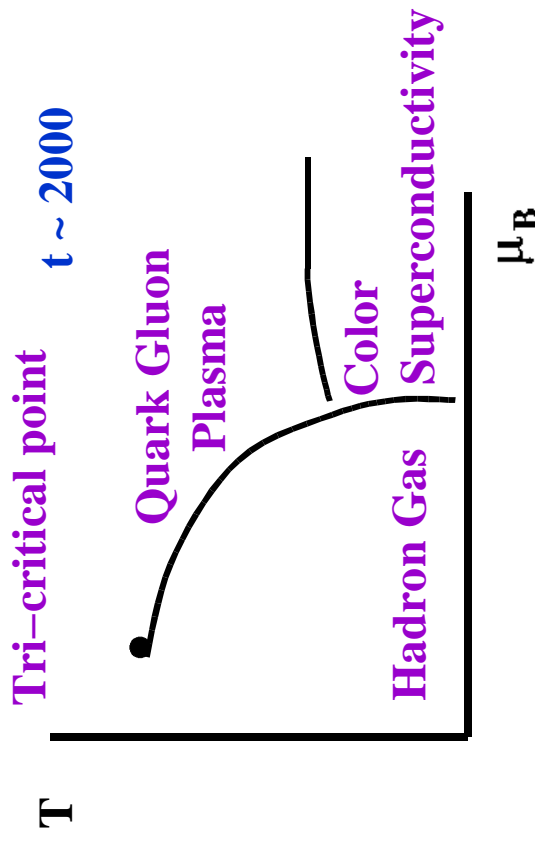
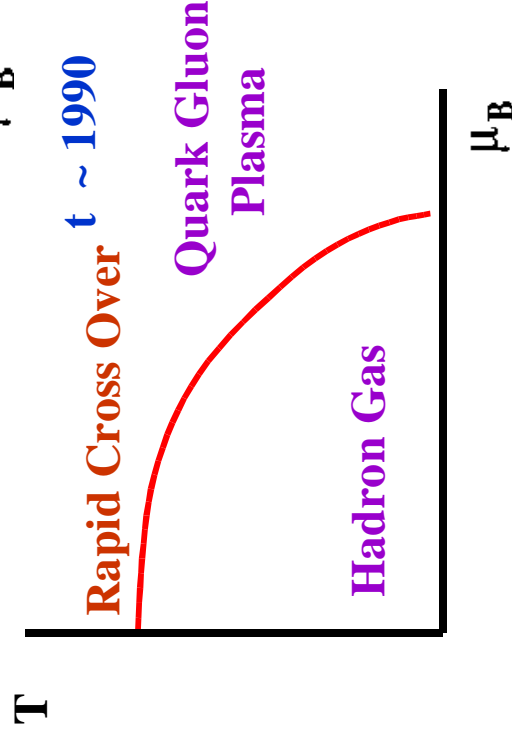


Critical Temperature

150 – 200 MeV

Critical Density

$\frac{1}{2} - 2$ Baryons/Fm³



Recent work stimulated by Wilczek and Rajagopal; Schaeffer and Shuryak on Color Superconductivity; also Stephanov, Son, Pisarski, Rischke.

What have lattice simulations shown?

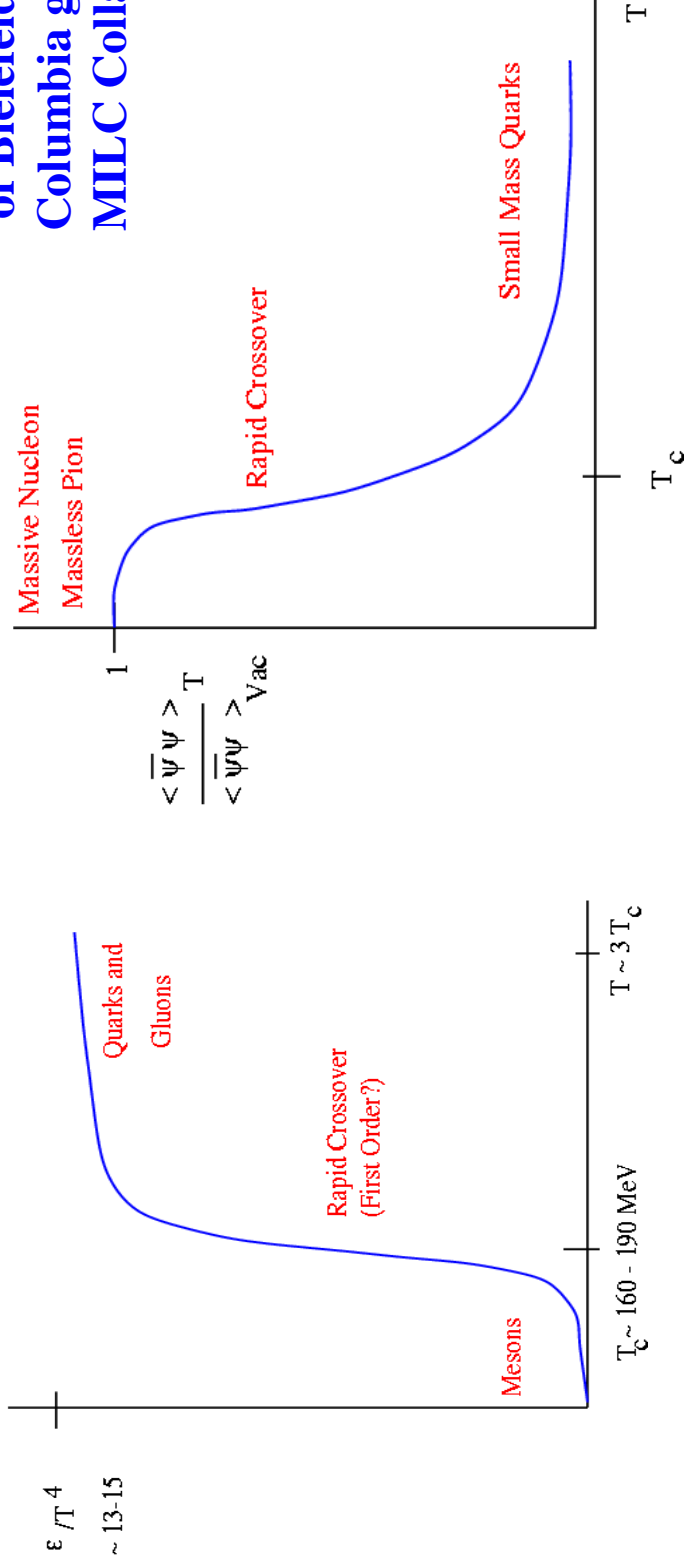
Chiral symmetry : $m_{\text{up}}, m_{\text{down}} \sim 0, M_{\text{nucleon}} \sim 1 \text{ GeV}$

How do particles get their mass?

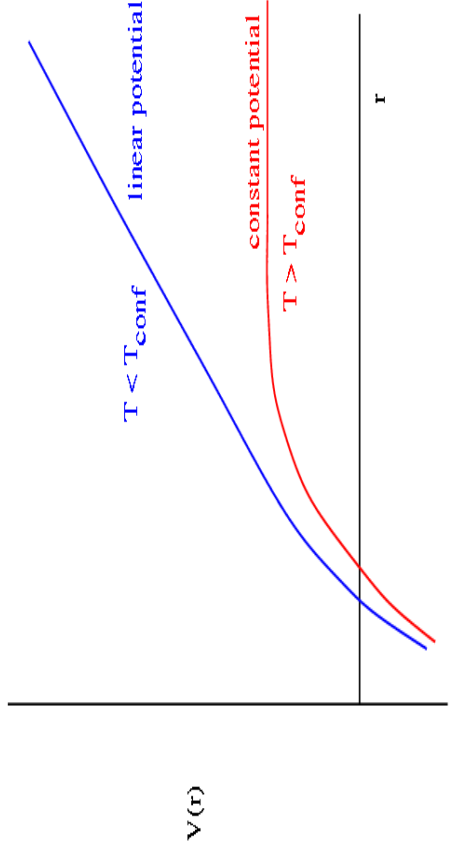
Why is the pion mass so small?

Is this related to confinement?

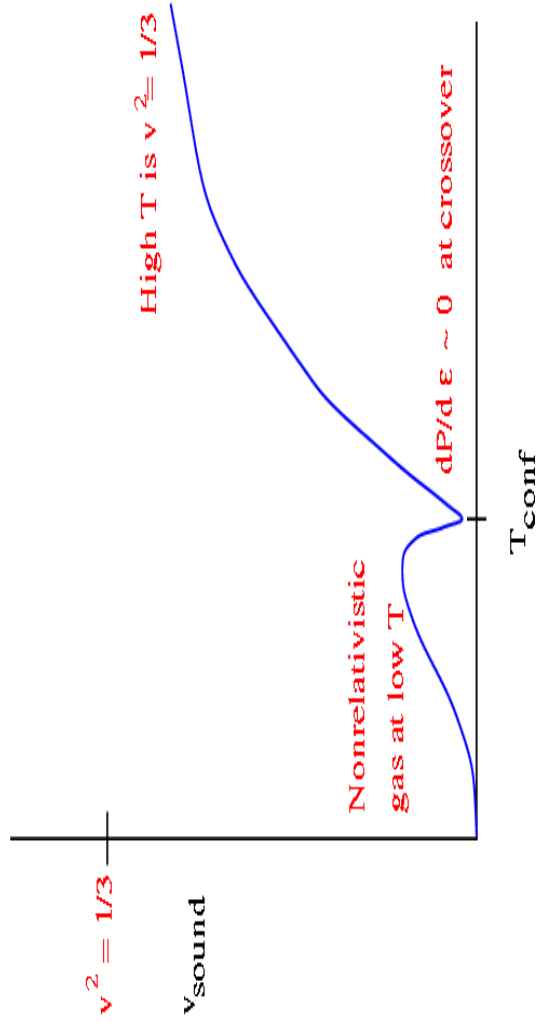
**Lattice simulations
of Bielefeld group,
Columbia group,
MILC Collaboration**



Is the confining force still linear at $T > T_{\text{dec}}$?



What is the equation of state? Sound velocity?



The Color Glass Condensate

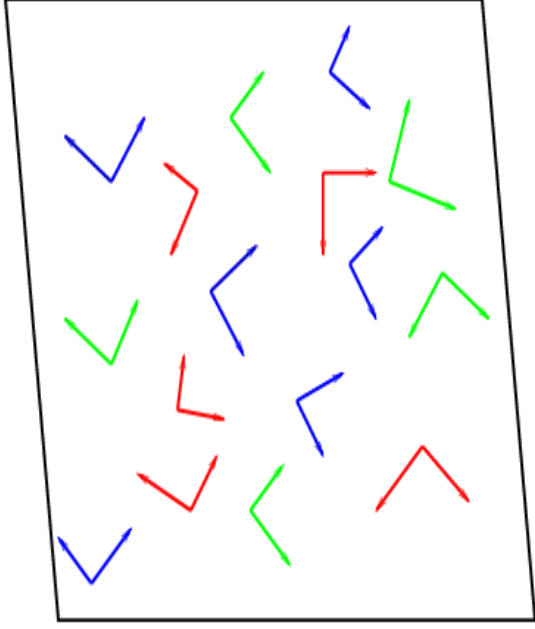
Hadron in frame where it has high momenta:

High momenta constituents generate low momentum wee partons.

Density of gluons per unit area becomes large.

Fields are random on thin sheet traveling near speed of light.

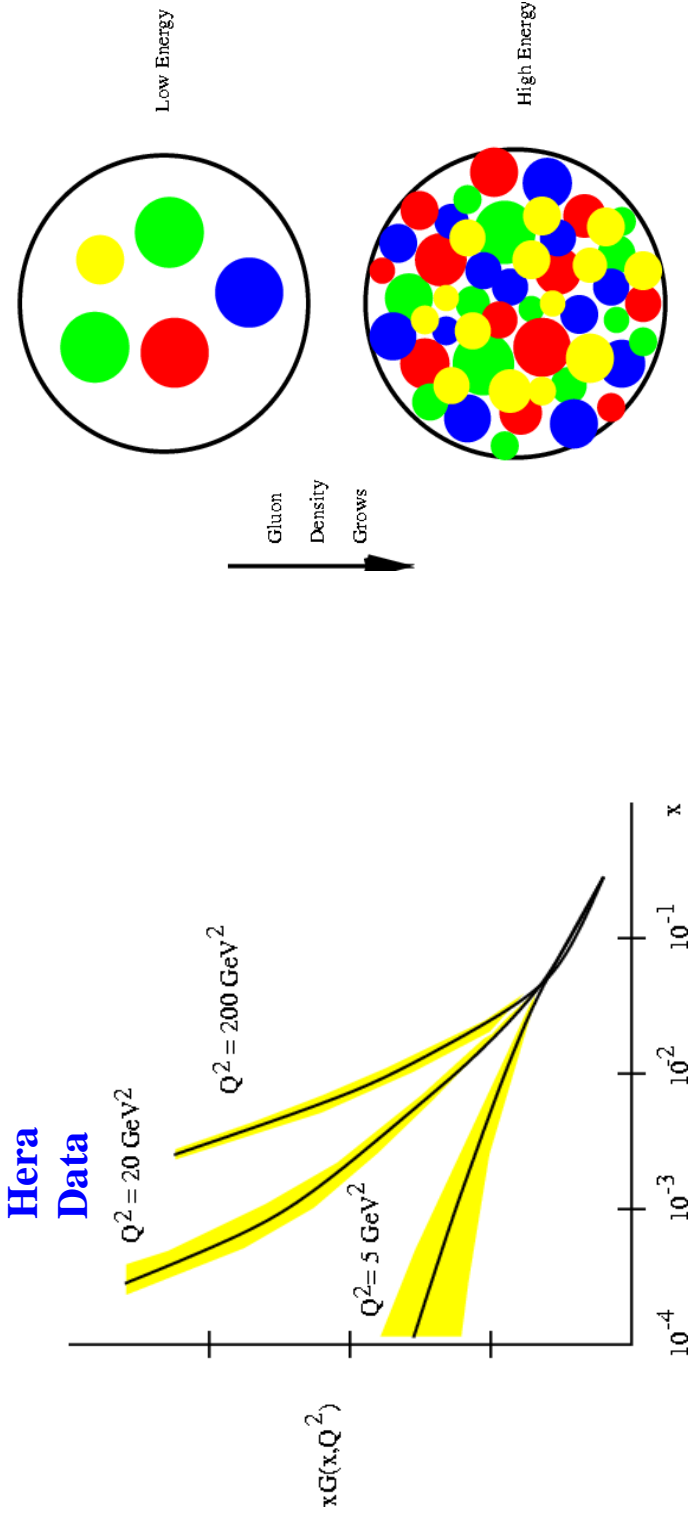
Universal high energy behavior for all hadrons.



↑
 $v \sim c$

Random Non-Abelian
Wiegacker-Williams
Fields

Color Glass Condensate and Saturation



Gribov, Levin, Ryskin;
Mueller; McLerran, Venugopalan

Gluon phase space density per unit area,

$$\rho = \frac{1}{\pi R^2} \frac{dN}{dp_T^2}$$

grows until

$$\rho \sim Q_{\text{sat}}^2 / \alpha_{\text{strong}}$$

$$Q_{\text{sat}} \gg \Lambda_{\text{QCD}} \Rightarrow \alpha_{\text{strong}} \ll 1$$

Color Glass Condensate

Minnesota Mob and
East Coast and
European Affiliates

Color: Made of colored gluons

Glass: Wee fields (low momentum) are produced by higher momentum constituents. Time scales are Lorentz contracted compared to natural time scales

Condensate: Gluon density as large as it can

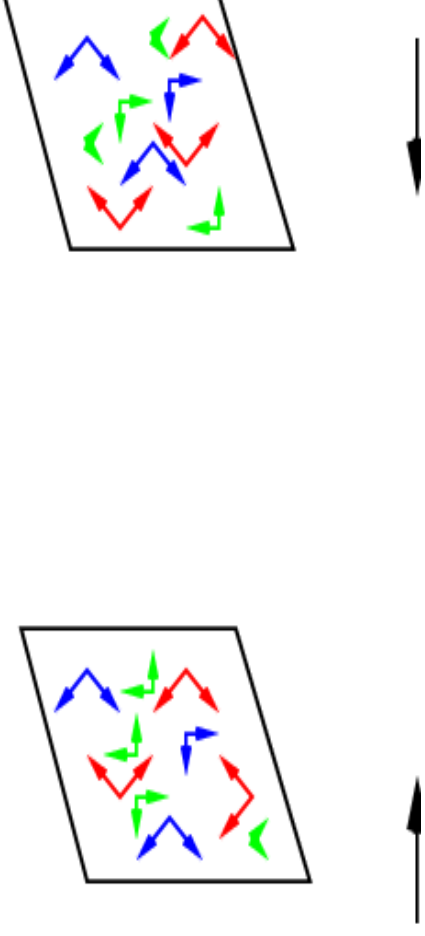
$$\rho \sim Q_{\text{sat}}^2 / \alpha_s$$

$1/\alpha_s$ is typical of Boson Condensation

If there are repulsive interactions, negative energy terms of order ρ become of order interaction terms, $\alpha \rho^2$

$$\text{so } \rho \sim 1/\alpha_s$$

Space Time Evolution of Ultrarelativistic Nuclear Collisions



McLerran, Kovner,
Weigert; Krasnitz,
Nara,
Venugopalan

$t < 0$

Color Glass Condensate in Each Nucleus

$0 < t < t_{\text{form}}$

Color Glass Condensate Melts Into Quark-Gluon Plasma

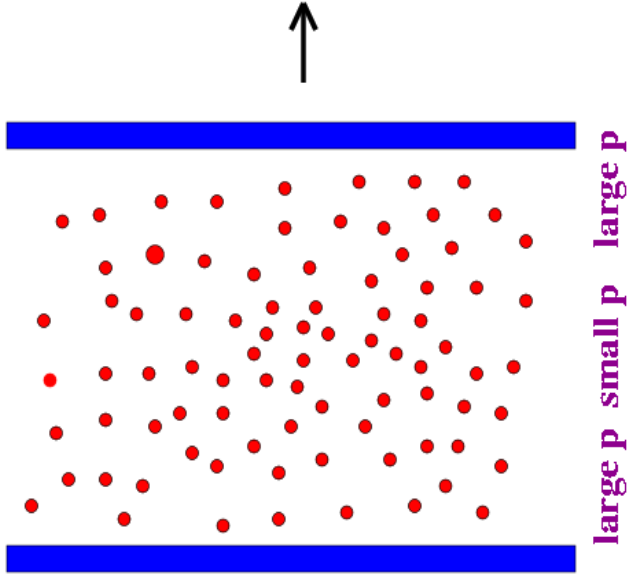
$t_{\text{form}} \sim 1/Q_{\text{sat}} \sim .3 \text{ Fm}/c$ at RHIC

$\epsilon_{\text{form}} \sim Q_{\text{sat}}^4 / \alpha_s \sim 50 - 100 \text{ GeV}/\text{Fm}^3$ at RHIC

Space Time Evolution of Ultrarelativistic

Nuclear Collisions

Bjorken



Matter is formed with correlation between momentum and position due to Lorentz time dilation of the formation time

Similar to Hubble expansion in Cosmology:
 $N/V \sim 1/t$

If not equilibrated, $E/N \sim \text{constant}$

If equilibrated, $E/N \sim T, T^3 \sim N/V \sim 1/t$
 or $T \sim t^{-1/3}$

$t_{\text{form}} \lesssim t \lesssim t_{\text{therm}}$

Matter thermalizes and Hubble Expands

$t_{\text{form}} \lesssim t \lesssim t_{\text{decoupling}}$

Matter expands as thermal system

$t_{\text{therm}} \sim .5 - 1 \text{ Fm}/c$

$t_{\text{decoupling}} \sim R/v_s \sim 10 \text{ Fm}/c$

Space Time Evolution in Ultrarelativistic Nuclear Collisions

A variety of intermediate time scales between thermalization and decoupling:

Quark Gluon Plasma until it begins hadronizing into a mixed phase of quarks gluons and hadrons

Mixed phase expands until it is entirely a hadron gas

Decoupling occurs at roughly the time the matter is entirely hadron gas, at RHIC energies, $t \sim 10 \text{ Fm}/c$

Can show the temperature always slower than

$$T \sim t^{-1/3} < \tau \sim t^{-1/3}$$

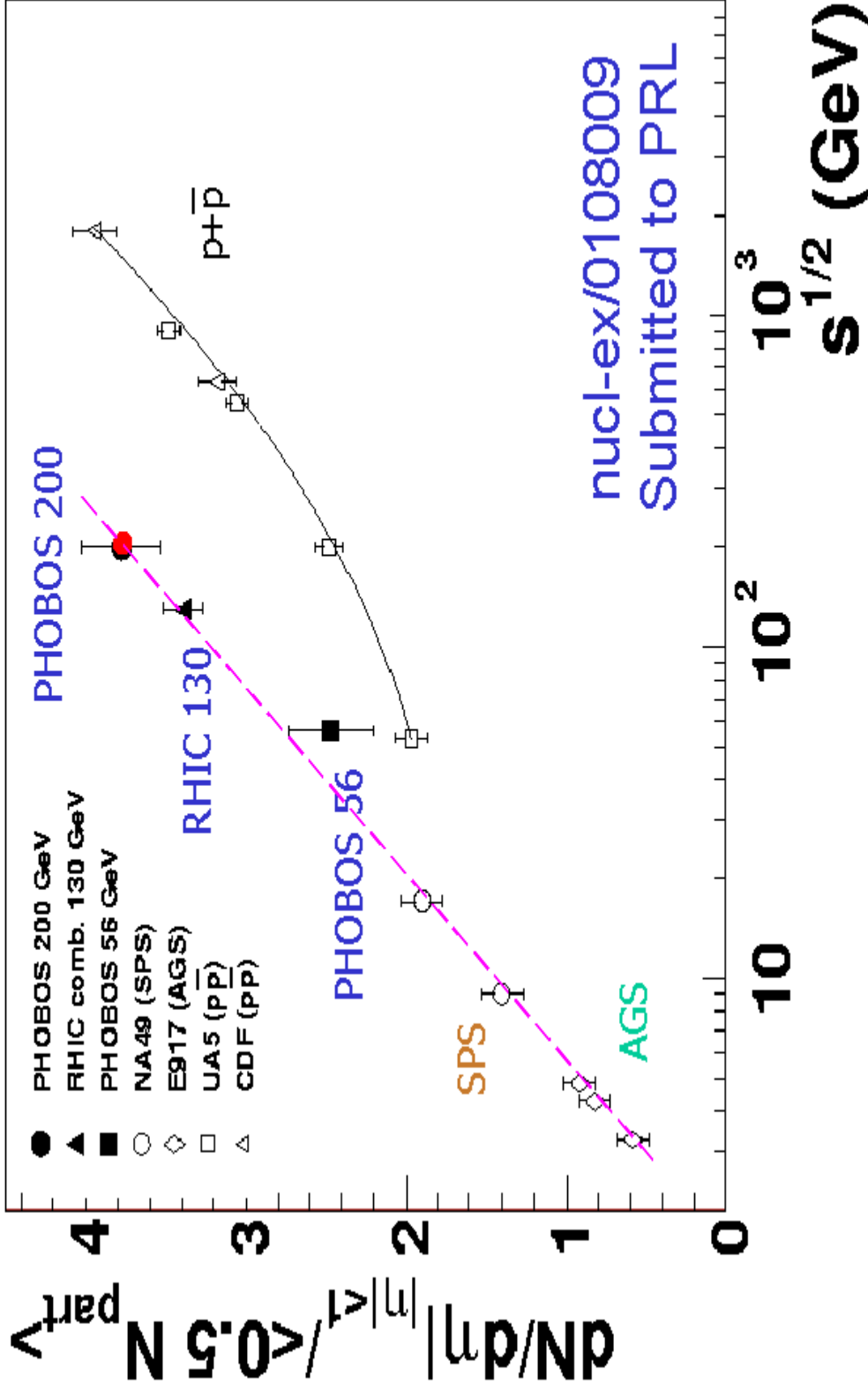
(so long as t is smaller than time at which expansion becomes

3 dimensional. System begins to decouple then)

What Have We Learned?

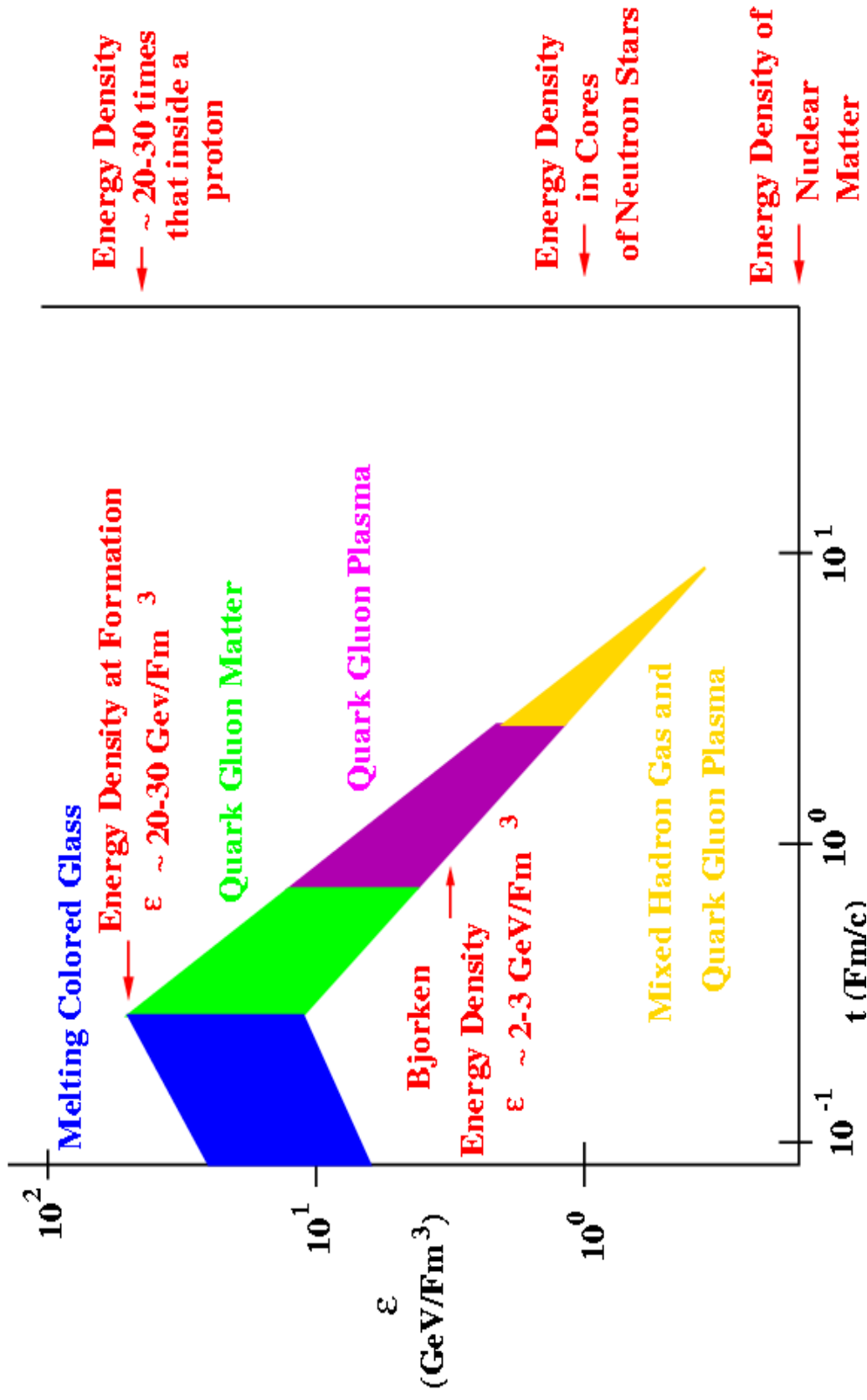
The Multiplicity as a Function of Energy

$dN_{ch}/dn_{|\eta|<1}$ vs Energy



What Have We Learned?

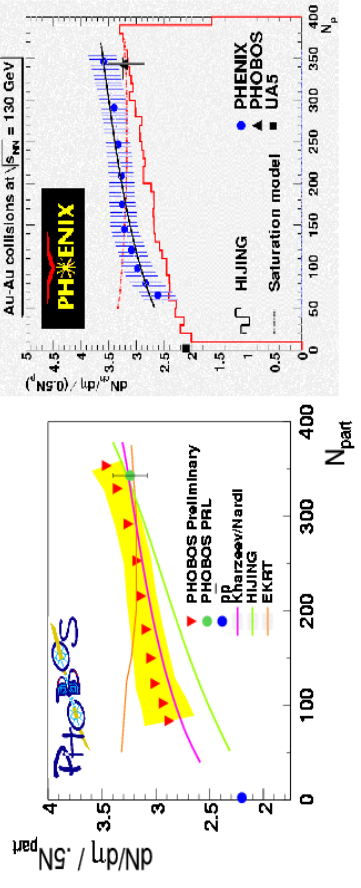
Bounds on energy density



Energy Density is Too Big for a Hadron Gas !

What Have We Learned?

Gross Properties of Multiplicities Consistent with Color Glass ! $dN/d\eta$ vs Centrality at $\eta=0$



Kharzeev and Nardi: $dN/dy \sim 1/\alpha_s$

$\propto (Q_{sat})$ depends on multiplicity per unit area

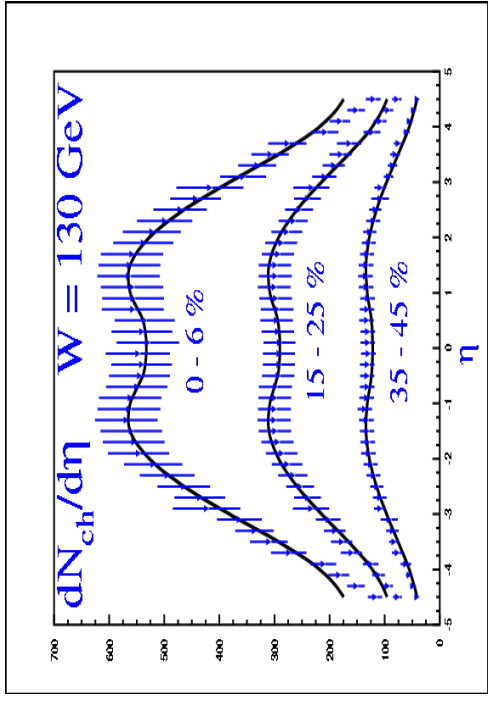
Kharzeev and Levin: Rapidity density as function of y can be computed in terms of Q_{sat}

Q_{sat} is the only scale !

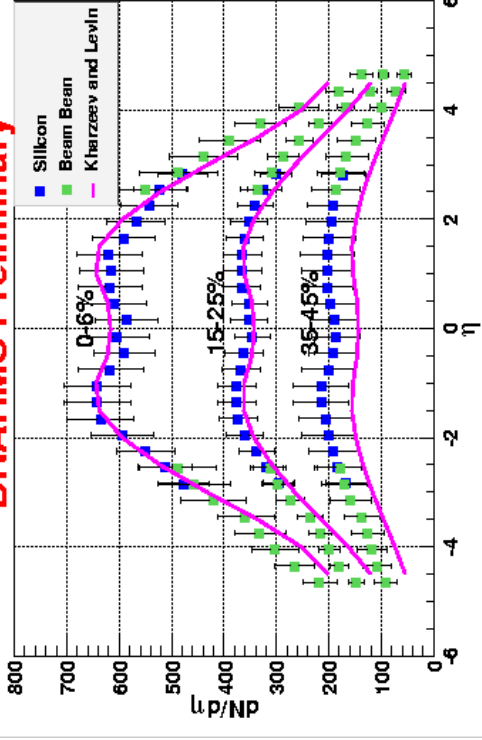
Peter Steinberg

BROOKHAVEN NATIONAL LABORATORY

PHOBOS



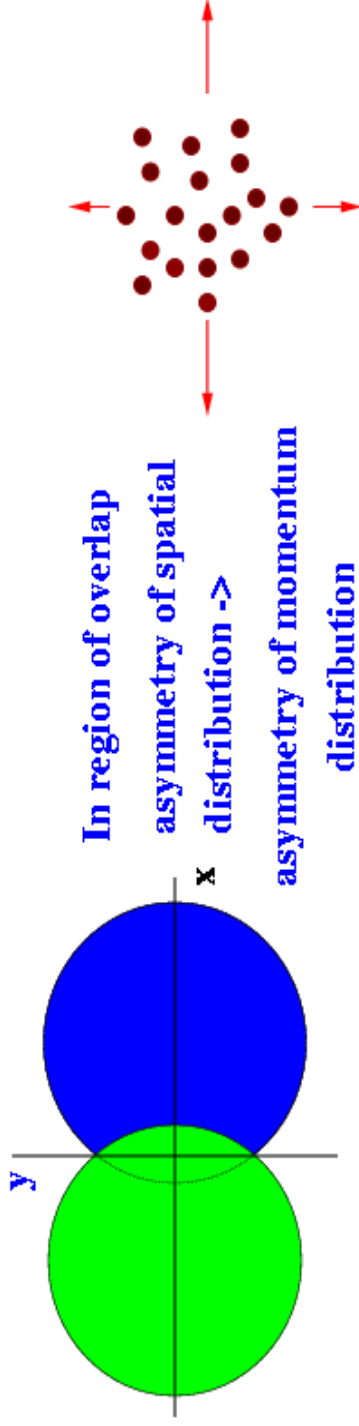
BRAHMS Preliminary



What Have We Learned?

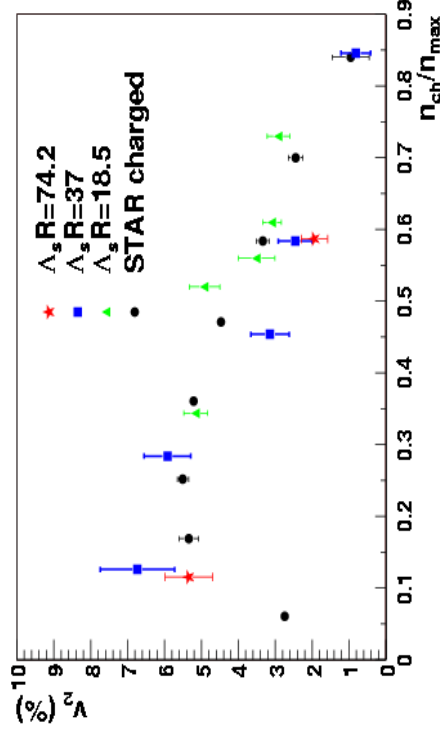
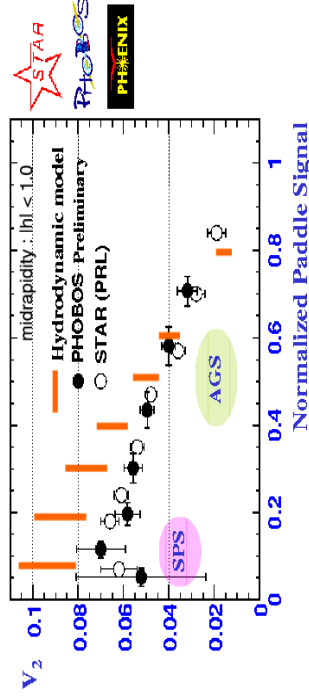
It is Matter and It Interacts Strongly!

Collective Flow and v2



$$V_2 = \langle \cos(2\phi) \rangle \quad \tan(\phi) = P_y/p_x$$

Centrality Dependence



Peter Steinberg

BROOKHAVEN
NATIONAL LABORATORY

Heinz et al;
Teaney, Shuryak

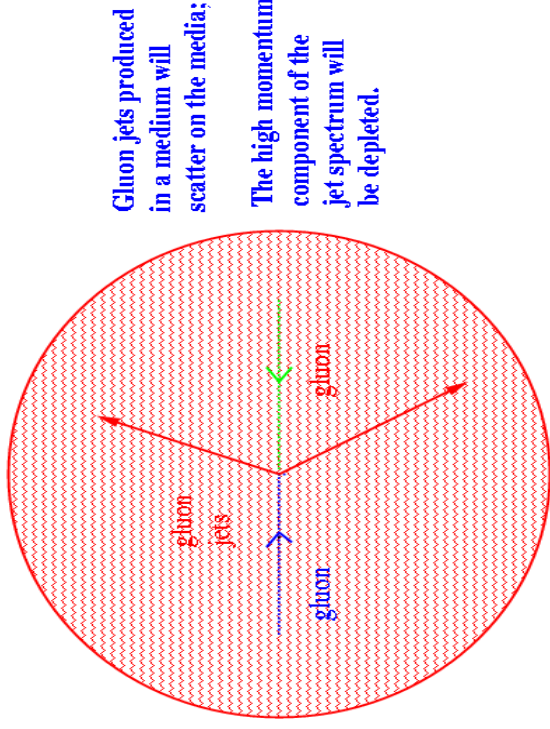
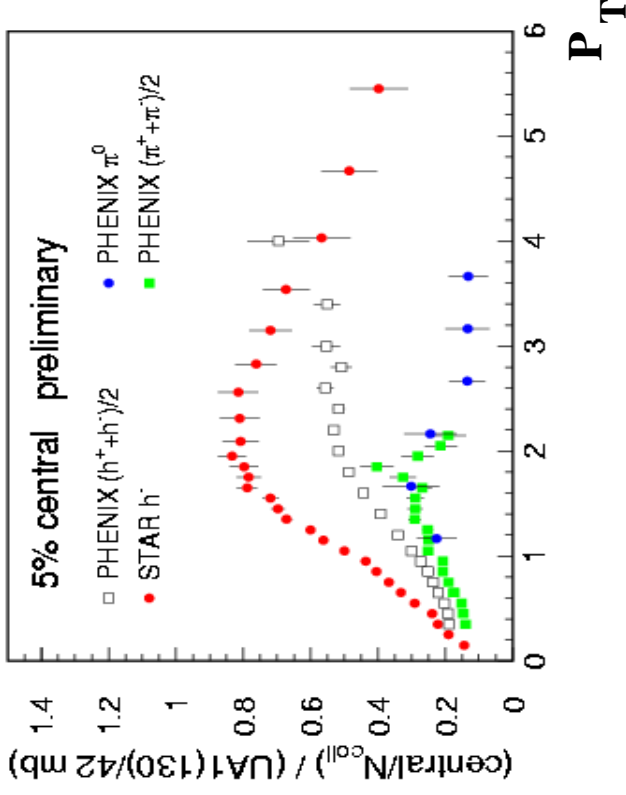
Nara, Krasnitz,
Venugopalan

Quark Gluon Plasma or Colored Glass?

(or both?)

What Would Do We Expect to Learn?

Does the matter equilibrate?



Gluon jets produced in a medium will scatter on the media;

The high momentum component of the jet spectrum will be depleted.

Problems with interpretation:

- 1) Large uncertainty in pp data
- 2) Systematic uncertainty in AA data
- 3) Nuclear modification of gluon distribution
- 4) Transverse momentum limited by statistics

Sytematics reduced by comparing data at same energy in same detector

Nuclear modification of distribution functions determined by pA (or eA)

Higher transverse momentum with higher accelerator luminosity

Gyulassy, Wang; Dokshitzer, Mueller, Baier, Schiff; Levai; Wiedemann

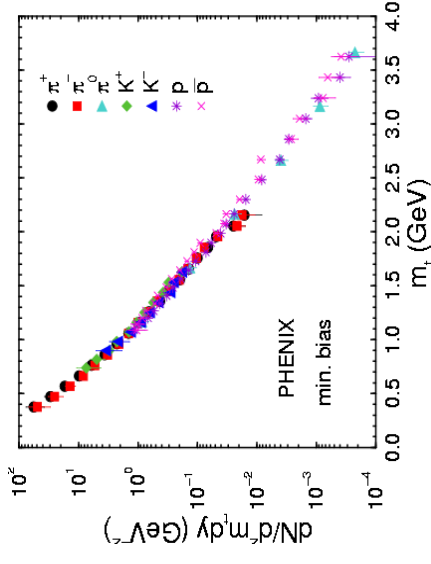
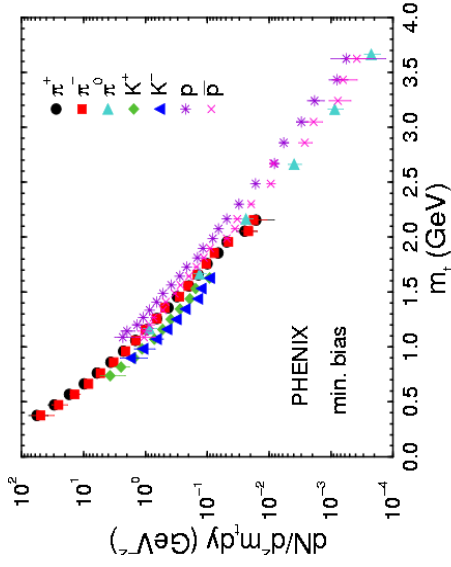
What Do We Expect to Learn?

Schaffner–Bielich et al

Is the pT Spectrum Modified Because of Color Glass or Quark Gluon Plasma?

M_T scaling: At fixed impact parameter, distributions are functions only of M_T up to an overall multiplicative constant dependent on particle species

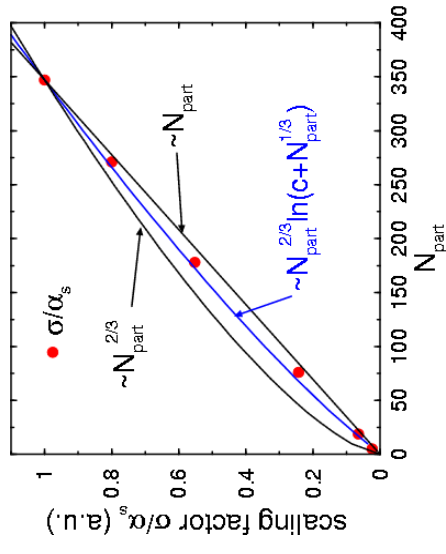
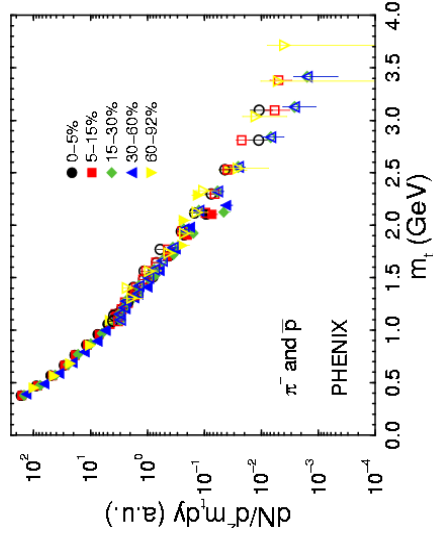
$$M_T^2 = p_T^2 + M^2$$



Color Glass Scaling:

$$\frac{dN}{d^2p_T} = \frac{\pi R^2}{\alpha_s} F(M_T/Q_{sat})$$

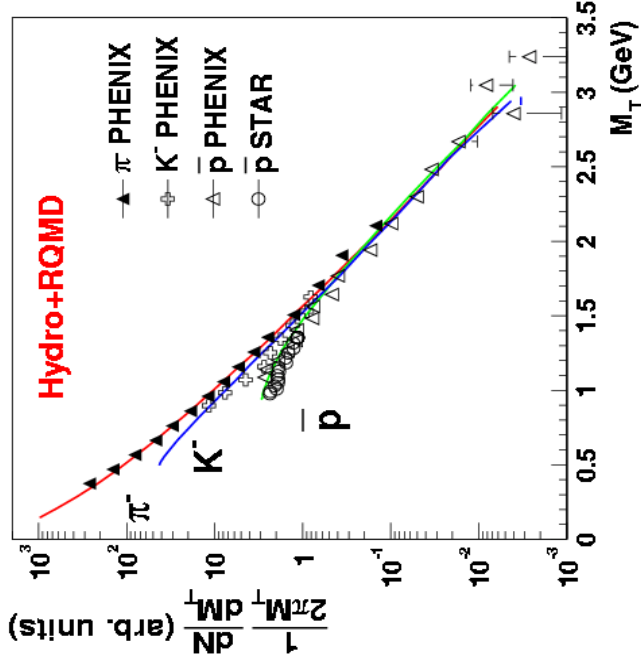
where $\alpha = \alpha(Q_{sat})$



Results are consistent with Color Glass picture!

What Do We Expect to Learn?

Also Consistent with Hydrodynamics of Quark Gluon Plasma !



For Hydrodynamics:

Shape of curve predicted in hydrodynamics.
Deviations from M_T scaling at $M_T \sim M$.
Deviation larger for more massive particles

For Color Glass:

Non-trivial relations for different centrality
Correctly generates dependence of strong coupling

Probably both contribute?

Heinz, Huovinen, Kolb;
Shuryak, Teaney; Hirono

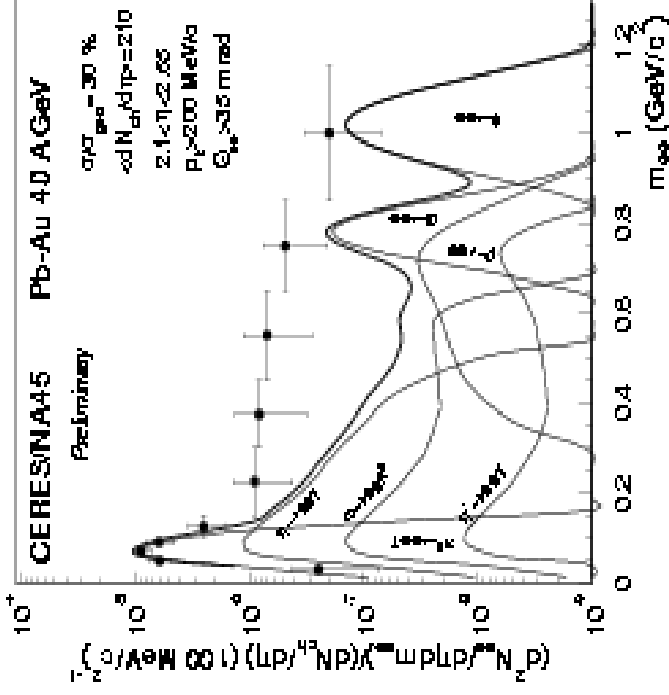
Srivastava

**If hydrodynamics is valid, can
determine equation of state!**

What Do We Hope to Learn?

Does deconfinement occur in high density matter?

Does chiral symmetry restoration change particle masses?



Spectrum of low mass dileptons:
Measured at CERN

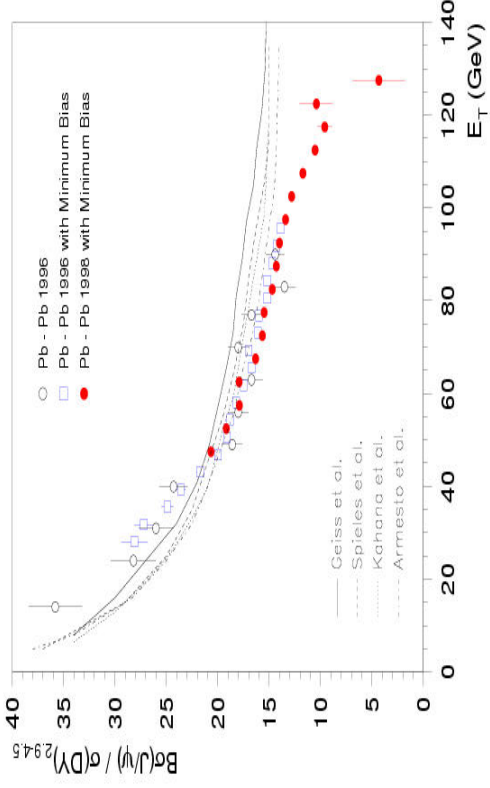
Probably resonance broadening
Rapp, Weise, Wambach

Difficult to measure low mass pairs
at RHIC due to backgrounds.

Very difficult, but very powerful technique.

What Do We Hope to Learn?

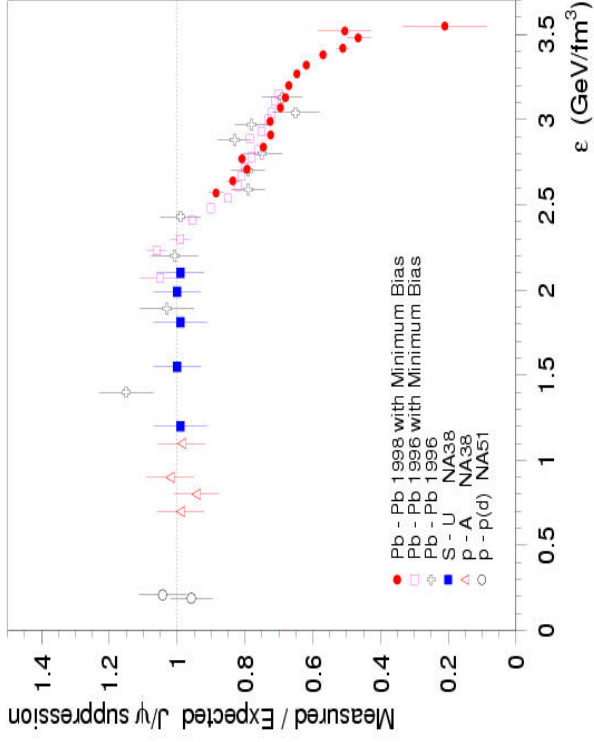
Melting of the J/Psi



Is J/ψ suppression due to hadronic scattering? Changes in gluon distribution function? Higher twist effects? Deconfinement?

At RHIC energies, does increased open charm production result in increased probability of final state recombination into charm.

Is the J/ψ enhanced at RHIC energies?



Need independent measurement of open charm

Must understand energy dependence.

Measurements in pA and pp essential!

Suppressors: Matsui, Satz Kharzeev
Blaizot, Ollitrault

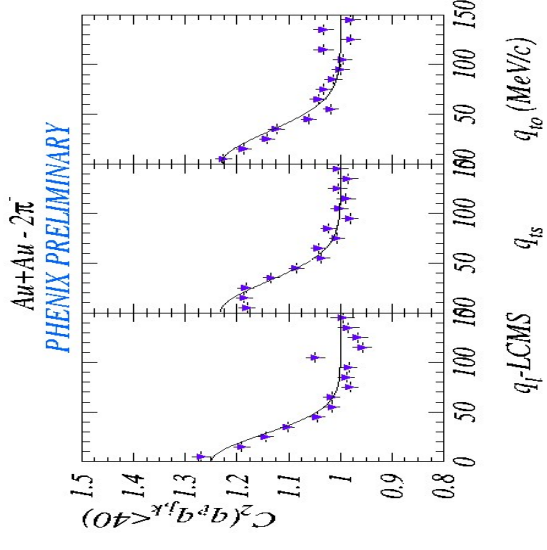
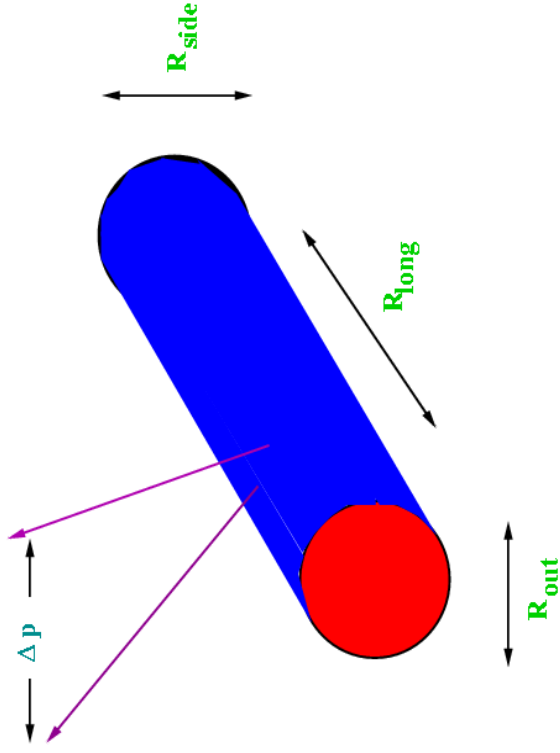
Twister: Qiu

Enhancers:

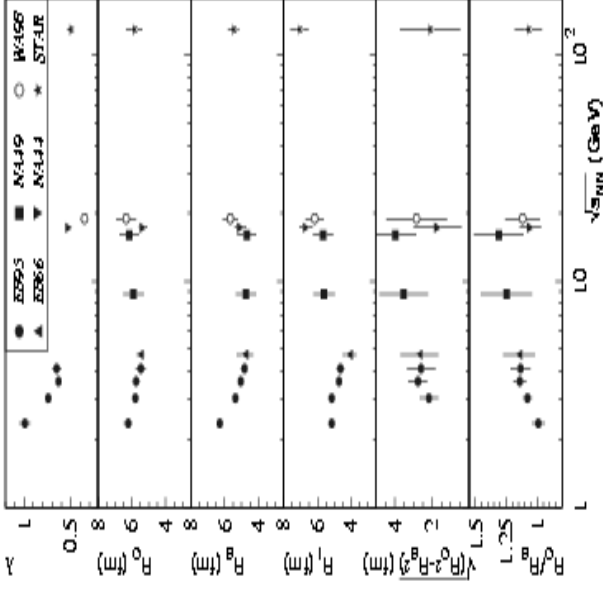
Rescatter: Capella
Rafelski, Thews,
Stachel, Braun–Munzinger,
Redlich, Gorenstein

What Do We Hope to Learn?

Lifetime and spatial extent of matter produced



- R_{side} = 4.42
(0.22) Fm
- R_{out} = 4.45
(0.22) Fm
- R_{long} = 5.28
(0.32) Fm



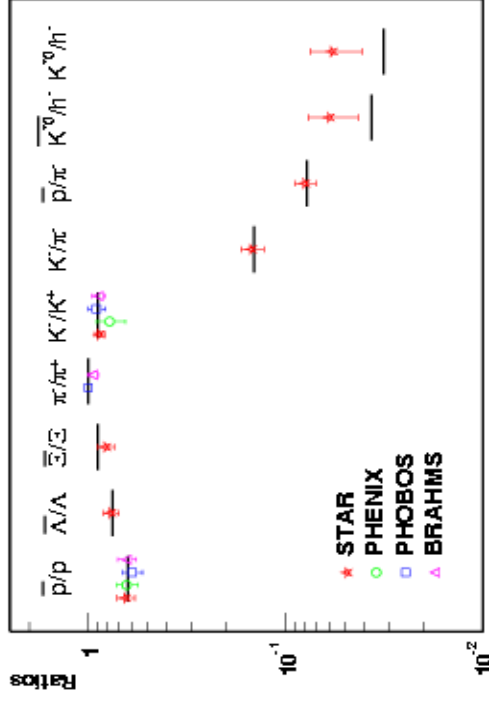
Difficult to interpret the ratio of R_0/R_s :
 Might be due to Color Glass initial conditions?
 Might be new phenomena associated with phase transition?

Do we really understand the space-time evolution ?

Soff, Bass, Dumitru; Kolb, Heinz; Teaney

What Do We Hope to Learn?

Flavor Composition of Quark–Gluon Plasma



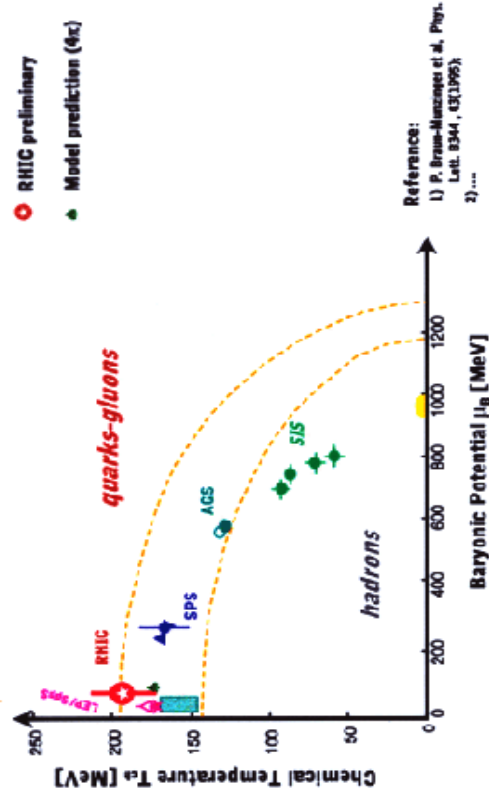
Can fit abundances with temperature and chemical potential for baryon number and strangeness.

It works too well! Works for electron–positron collisions!

Are we not understanding something fundamental and universal about hadron collisions?

Is strangeness abundance a signal for quark–gluon plasma?

Mueller, Rafelski; Cleymans, Redlich; Braun–Munzinger, Stachel; Gorenstein Gazdzicki



Summary:

What are we trying to understand:

New forms of matter:

Color Glass Condensate and Quark Gluon Plasma
(Color Glass at early times, Plasma later)

What have we already learned:

Matter has been produced at energy densities so high it
can only be made from quarks and gluons

The matter is strongly interacting

Multiplicity distributions consistent with **Colored Glass**

Flow and pT distributions consistent with both
Colored Glass and Quark–Gluon Plasma

What do we expect to learn:

By measuring low pT heavy hadrons, resolve differences between various theory descriptions.

By measuring high pT hadrons in AA and pA collisions determine whether jet quenching occurs, and to what degree thermalization occurs.

**If thermalization is established, then can measure properties of
Equation of State**

What do we hope to learn:

Through lepton pair measurements,
whether or not resonances shifted in mass, broadened or melted.

What the HBT determined sizes tell us about the space–time pictures.

What flavor abundances tell us about properties of high density matter

We must turn to the problem of determining the properties of the high density matter made in heavy ion collisions

We are learning what are the useful theoretical and experimental tools

Will need experiments on AA and pA to resolve:

Initial state effects: Color Glass Condensate

Evolution of the matter: Quark Gluon Plasma
(eA for precise determination of Color Glass)

