

Analogous to the non-relativistic force relation

$$\vec{f}_{\text{non-rel}} = m \vec{a} = \frac{d}{dt} (m \vec{v}) = \frac{d}{dt} (\vec{p}) \quad (\text{Non-relativistic})$$

we introduce the four-force

$$F = \frac{dP}{d\tau} \left(= \frac{d}{d\tau} (m_0 u) \right) = \frac{d}{d\tau} \left(\frac{E}{c}, \vec{P} \right)$$

(Like Rindler I here let capital letters without indices denote contravariant four-vectors.)

If m_0 is constant, generally m_0 may change as energy ~ mass

where P is the contravariant four-momentum.

Since P is a four-vector and τ is a scalar F is a four-vector.

Using

$$\frac{d}{d\tau} = \frac{dt}{d\tau} \frac{d}{dt} = \gamma(\vec{v}) \frac{d}{dt}$$

we get

$$P = \gamma(\vec{v}) m_0 (c, \vec{v})$$

introduce $\vec{f} = \frac{d\vec{P}}{dt}$

$$F = \gamma(\vec{v}) \frac{d}{d\tau} P = \gamma(\vec{v}) \frac{d}{dt} (\gamma(\vec{v}) m_0 c, \vec{P}) = \gamma(\vec{v}) \left(\frac{1}{c} \frac{dE}{dt}, \vec{f} \right)$$

Consider

$$F \cdot U = \gamma(\vec{v}) \left(\frac{1}{c} \frac{dE}{dt}, \vec{f} \right) \cdot \gamma(\vec{v}) (c, \vec{v})$$

$$= \gamma(\vec{v})^2 \left(\frac{dE}{dt} - \vec{f} \cdot \vec{v} \right) = c^2 \frac{dm_0}{d\tau}$$

This is a scalar => same in all systems
=> consider in rest system in which $E = m_0 c^2, \vec{v} = 0, \gamma = 1$

From this we see

$$\frac{d}{d\tau} m_0 = 0 \Leftrightarrow \frac{d}{dt} m_0 = 0 \Leftrightarrow F \cdot U = 0$$

(A force for which m does not change is called pure in Rindler)

i.e., the (rest) mass is conserved iff $F \cdot U = 0$

In fact, as remarked before, Maxwell's equations are already invariant under Lorentz transformations. Today we will assume we do not know them and ask the question:

What form can a relativistic electromagnetic force possibly have?

Assume:

- It preserves the rest mass => $F \cdot U = 0$

(expected of electromagnetic force, but not in quantum field theories, where particles are created)

- It is proportional to the charge

1) Assume (incorrectly) that the four-force $F = (F^0, F^1, F^2, F^3)$ does not depend on the velocity, but then it cannot be a four-vector, since it cannot transform like one. (The force can thus only be velocity independent in one frame.)

2) Next assume $F \propto U$ but then $F \cdot U \propto U \cdot U = c^2 \neq 0$
 => The force cannot preserve (rest) mass.

3) The third simplest guess is that the force depends linearly on the four-velocity:

$$F_\mu (\equiv g_{\mu\nu} F^\nu) = \frac{q}{c} F_{\mu\nu} U^\nu$$

covariant four-vector factor c for convenience four-velocity, contravariant four-vector

(The tensor $F_{\mu\nu}$ is called $F_{\mu\nu}$ almost everywhere, but Rindler calls it $E_{\mu\nu}$)

must transform as a rank $\begin{pmatrix} 0 \\ 2 \end{pmatrix}$ tensor

where $F_{\mu\nu}$ is the electromagnetic field-tensor.

Next we require that the rest mass doesn't change $\Rightarrow F \cdot U = 0$

L5:3

$$F \cdot U = F_{\mu} U^{\mu} = \frac{q}{c} F_{\mu\nu} \underbrace{U^{\nu} U^{\mu}} \stackrel{!}{=} 0$$

symmetric in $\mu \leftrightarrow \nu$

so we see that if $F_{\mu\nu}$ is antisymmetric in $\mu \leftrightarrow \nu$ i.e., $F_{\mu\nu} = -F_{\nu\mu}$

then $F \cdot U = 0 \Rightarrow$ rest mass conserved.

Ex: The sum contains the 13 and 31-terms:

$$F_{13} U^1 U^3 + F_{31} U^3 U^1 = \underbrace{(F_{13} + F_{31})}_{0 \text{ if } F_{31} = -F_{13}} U^1 U^3$$

We have assumed that the fields affect the charges in proportion to U ,

let's assume that the charges affect the fields in proportion to U

$$\Rightarrow \partial_{\mu} F^{\mu\nu} \equiv F^{\mu\nu}_{, \mu} \stackrel{\text{def of } , \mu}{=} k \rho_0 U^{\nu} \equiv k J^{\nu} \quad (*) \quad \text{(We can raise and lower indices as we want with } g^{\mu\nu} \text{)}$$

where we differentiate w.r.t. all indices standing after ,

• ρ_0 is the proper charge density, the density in the rest system

• J^{μ} is a four-vector defined by this equation, the four-current

$$\text{From } k J^{\nu}_{, \nu} \stackrel{(*)}{=} F^{\mu\nu}_{, \mu\nu} = -F^{\nu\mu}_{, \mu\nu} = -\underbrace{F^{\nu\mu}_{, \nu\mu}} = -k J^{\nu}_{, \nu}$$

differ only by dummy indices

we conclude

$$\boxed{J^{\nu}_{, \nu} = 0}$$

the continuity equation

diagonal entries are 0

Cf.

$$\frac{\partial \rho}{\partial t} + \nabla \cdot \vec{j} = 0$$

Note that $F^{\mu\nu}$ has $\frac{4^2 - 4}{2} = 6$ components

antisymmetry

but we only have 4 field equations (*)

\Rightarrow Not enough information to constrain $F^{\mu\nu}$?

Compare this to the gravitational force, having three components, but being the derivative of one scalar field. (One d.o.f. giving forces in 3 directions.)

The situation here can be solved similarly, by letting $F^{\mu\nu}$ be the (antisymmetrized) derivative of a four-potential Φ

$$F^{\mu\nu} = \partial^\mu \Phi^\nu - \partial^\nu \Phi^\mu = \Phi^{\nu,\mu} - \Phi^{\mu,\nu}$$

($F_{\mu\nu}$ has to be antisymmetric so we cannot just take $F^{\mu\nu} = \partial^\mu \Phi^\nu$)

From the field equations (*) we then have

$$F^{\mu\nu}{}_{,\mu} = \Phi^{\nu,\mu}{}_{,\mu} - \Phi^{\mu,\nu}{}_{,\mu} = kJ^\nu$$

(Φ^μ is most often called A^μ)

Given boundary conditions on Φ , these four equations are enough to constrain Φ which in turn gives $F^{\mu\nu}$ (Only 4 d.o.f. in Φ)

However, different potentials $\Phi, \tilde{\Phi}$ can give rise to the same fields $F^{\mu\nu}$

Let $\tilde{\Phi} = \Phi - \psi$ then

$$\begin{aligned} F^{\mu\nu} \rightarrow \tilde{F}^{\mu\nu} &= \partial^\mu \tilde{\Phi}^\nu - \partial^\nu \tilde{\Phi}^\mu \\ &= \partial^\mu (\Phi^\nu - \psi^\nu) - \partial^\nu (\Phi^\mu - \psi^\mu) \\ &= \partial^\mu \Phi^\nu - \partial^\nu \Phi^\mu - \partial^\mu \psi^\nu + \partial^\nu \psi^\mu \\ &= F^{\mu\nu} - \partial^\mu \psi^\nu + \partial^\nu \psi^\mu \end{aligned}$$

So if $-\partial^\mu \psi^\nu + \partial^\nu \psi^\mu = 0$ we get the same fields $F^{\mu\nu}$

but this is the case if $\psi^\mu = \partial^\mu \varphi$ for some scalar φ

since then $-\partial^\mu \psi^\nu + \partial^\nu \psi^\mu = -\partial^\mu \partial^\nu \varphi + \partial^\nu \partial^\mu \varphi = 0$

derivatives commute

\Rightarrow We get the same field if Φ is changed by a total derivative $\partial^\mu \varphi$,

we have gauge invariance.

In fact we have already seen this in the analytical mechanics part where

our gauge invariance condition was

This ' should not be confused with a Lorentz transformed field

$$\tilde{\Phi}^\mu = \Phi^\mu - \partial^\mu \varphi$$

(AM L6:6-7)
$$\left\{ \begin{array}{l} V \rightarrow V' = V - \frac{\partial \Lambda}{\partial t} = V - \frac{\partial}{\partial x^0} c \Lambda \Leftrightarrow \tilde{\Phi}^0 = \Phi^0 - \frac{\partial}{\partial x_0} \varphi \quad \text{0-component} \\ \vec{A} \rightarrow \vec{A}' = \vec{A} + \nabla \Lambda \Leftrightarrow \frac{\tilde{\Phi}^i}{c} = \frac{\Phi^i}{c} + \frac{\partial \varphi}{\partial x^i} \quad \text{spatial components} \end{array} \right.$$

$\Phi^0 = V, \quad \varphi = c \Lambda, \quad \vec{\Phi} = c \vec{A}$

Divide with c to get Λ : $\vec{\Phi} = c \vec{A}$

$$\Phi = (V, c \vec{A})$$

There the E and B-fields were given by

$$\begin{aligned} \vec{E} &= -\nabla V - \frac{\partial \vec{A}}{\partial t} & \vec{B} &= \nabla \times \vec{A} = \nabla \times \frac{\vec{\Phi}}{c} & \nabla &= \left(\frac{\partial}{\partial x}, \frac{\partial}{\partial y}, \frac{\partial}{\partial z} \right) \\ &= -\nabla V - \frac{\partial}{\partial t} \frac{\vec{\Phi}}{c} & & & &= \left(\frac{\partial}{\partial x^1}, \frac{\partial}{\partial x^2}, \frac{\partial}{\partial x^3} \right) \\ &= -\nabla \Phi^0 - \partial_0 \frac{\vec{\Phi}}{c} & & & &= (\partial_1, \partial_2, \partial_3) \end{aligned}$$

In component form:

$$\begin{aligned} E^i &= -\partial_i \Phi^0 - \partial_0 \Phi^i \\ &= \partial^i \Phi^0 - \partial^0 \Phi^i \\ &= F^{i0} \end{aligned}$$

$$\begin{aligned} c B^1 &= \partial^3 \Phi^2 - \partial^2 \Phi^3 = F^{32} \\ c B^2 &= \partial^1 \Phi^3 - \partial^3 \Phi^1 = F^{13} \\ c B^3 &= \partial^2 \Phi^1 - \partial^1 \Phi^2 = F^{21} \end{aligned}$$

$$B^i = \frac{1}{c} \epsilon^{ijk} \partial_j \Phi^k$$

$$c B^i = -\epsilon^{ijk} \partial_j \Phi^k$$

$$\left\{ \begin{array}{l} \epsilon^{123} = \epsilon^{231} = \epsilon^{312} = 1 \\ \epsilon^{213} = \epsilon^{132} = \epsilon^{321} = -1 \\ \text{all others are 0} \end{array} \right.$$

(Since we are equating E and B fields with 3 components to tensors, it is not possible to use a manifestly covariant notation)

We thus see that

$$F^{\mu\nu} = \begin{bmatrix} 0 & F^{01} & F^{02} & F^{03} \\ -F^{01} & 0 & F^{12} & F^{13} \\ -F^{02} & -F^{12} & 0 & F^{23} \\ -F^{03} & -F^{13} & -F^{23} & 0 \end{bmatrix} = \begin{bmatrix} 0 & -E^1 & -E^2 & -E^3 \\ E^1 & 0 & -cB^3 & cB^2 \\ E^2 & cB^3 & 0 & -cB^1 \\ E^3 & -cB^2 & cB^1 & 0 \end{bmatrix}$$

We conclude that the \vec{E} & \vec{B} fields are part of a rank $\begin{pmatrix} 2 \\ 0 \end{pmatrix}$ tensor.

\Rightarrow We know how the \vec{E} & \vec{B} fields transform!

In terms of the \vec{E} & \vec{B} fields, the equation

$$F^\mu = \frac{q}{c} F^{\mu\nu} U_\nu$$

gives

$$F^1 = \frac{q}{c} F^{1\nu} U_\nu = \frac{q}{c} (F^{10} U_0 + 0 + F^{12} U_2 + F^{13} U_3)$$

$$= \frac{q}{c} (E^1 U^0 - cB^3 U_2 + cB^2 U_3)$$

$$U = \gamma(c, \vec{v}) \quad \Rightarrow \quad \frac{q}{c} \gamma (E^1 c - cB^3(-v^2) + cB^2(-v^3))$$

From

$$F = \gamma(\vec{v}) \left(\frac{1}{c} \frac{d\vec{E}}{dt}, \vec{f} \right) \quad \Rightarrow \quad \vec{f} = \frac{1}{\gamma(\vec{v})} \vec{F}$$

we get for the three-force

$$\Rightarrow \vec{f} = \frac{\vec{F}}{\gamma} = q (E^1 + B^3 v_y - B^2 v_x)$$

but this is nothing but the first component of the Lorentz force!

$$(\vec{f}) = \left(q (\vec{E} + \vec{v} \times \vec{B}) \right)' = q (E_x + v_y B_z - v_z B_y)$$

So we have found the correct electric and magnetic fields.

From the field equations $\partial_\mu F^{\mu\nu} = k \rho_0 U^\nu = k J^\nu$ (*) we get:

(The below two equations determine how E and B are affected by charges.)

- $\nu = 0$

$$\partial_0 F^{00} + \partial_1 F^{10} + \partial_2 F^{20} + \partial_3 F^{30}$$

$$= 0 + \partial_1(E^1) + \partial_2(E^2) + \partial_3(E^3) = k \rho_0 U^0 \Leftrightarrow \nabla \cdot \vec{E} \stackrel{U^0 = \gamma c}{=} k \rho_0 \gamma c = k \rho c$$

- $\nu = 1$

$$\partial_0 F^{01} + \partial_1 F^{11} + \partial_2 F^{21} + \partial_3 F^{31} = k J^1$$

$$-\partial_0 E^1 + 0 + \partial_2 c B^3 + \partial_3 (-c B^2) = k J^1 \quad \nabla \times \vec{B} = \frac{k \vec{J}}{c} + \frac{1}{c^2} \frac{\partial \vec{E}}{\partial t}$$

$$\frac{1}{c} \frac{\partial}{\partial t} = \frac{1}{\epsilon_0 c^2} \vec{J} + \frac{1}{c^2} \frac{\partial \vec{E}}{\partial t}$$

i.e., two of Maxwell's equations, for the others consider:

- $\vec{B} = \nabla \times \frac{\vec{\Phi}}{c}$

$$\nabla \cdot \vec{B} = \nabla \cdot (\nabla \times \frac{\vec{\Phi}}{c})$$

(Part of Rindler eq. 38.5)

$$= \frac{1}{c} \partial_i \epsilon^{ijk} (\partial_j \Phi^k - \partial_k \Phi^j)$$

$$= \frac{1}{c} \epsilon^{ijk} (\partial_i \partial_j \Phi^k - \partial_i \partial_k \Phi^j) = 0$$

antisym. in i,j
antisym. in i,k
sym. in i,j
sym. in i,k

(The below two equations are consequences of the fact that the field is determined by a four-potential.)

(There cannot be any magnetic monopoles if electromagnetism is formulated like this)

- $(\nabla \times \vec{E})^1 = \partial_2 \vec{E}^3 - \partial_3 \vec{E}^2$

(Part of Rindler eq. 38.5)

$$= \partial_2 F^{30} - \partial_3 F^{20}$$

$$= (\cancel{\partial_2 \partial^3 \Phi^0} - \partial_2 \partial^0 \Phi^3) - (\cancel{\partial_3 \partial^2 \Phi^0} - \partial_3 \partial^0 \Phi^2)$$

$$= \partial^0 (-\partial_2 \Phi^3 + \partial_3 \Phi^2) = \partial^0 F^{23} = \frac{1}{c} \frac{\partial}{\partial t} (-c B^1) = -\frac{\partial}{\partial t} B^1$$

Similarly for 2nd and 3rd component $\Rightarrow \nabla \times \vec{E} = -\frac{\partial}{\partial t} \vec{B}$

Recall:

$$F^{\mu\nu} = \partial^\mu \Phi^\nu - \partial^\nu \Phi^\mu$$

\Rightarrow We have recovered Maxwell's equations!

Since we can change Φ by a total derivative without changing our \vec{E} & \vec{B} fields, we can choose gauge to make our calculations particularly simple.

One option is to let

$$\underline{\partial_\mu \Phi^\mu = 0.} \quad \underline{\text{Lorentz gauge condition}}$$

To see that we can fix this we note that if we choose φ s.t.

$$\partial_\mu \tilde{\Phi}^\mu + \partial_\mu \partial^\mu \varphi = 0$$

then, recalling $\Phi = \tilde{\Phi} + \partial^\mu \varphi$ we get

$$\underline{\partial_\mu \Phi^\mu} = \partial_\mu \tilde{\Phi}^\mu + \partial_\mu \partial^\mu \varphi = \underline{0} \quad \text{ok}$$

The field equations then give

$$\begin{aligned} F^{\mu\nu}{}_{,\mu} &= \Phi^{\nu,\mu}{}_{,\mu} - \Phi^{\mu,\nu}{}_{,\mu} \quad (\stackrel{(*)}{=} k J^\nu) \\ &= \partial_\mu \partial^\mu \Phi^\nu - \partial_\mu \partial^\nu \Phi^\mu \\ &= \square \Phi^\nu \end{aligned}$$

I.e. we can always choose gauge s.t. $\Rightarrow \square \Phi^\nu = k J^\nu$

In particular we note that if $J^\nu = 0$

$$\square \Phi^\nu = 0$$

For $F^{\mu\nu}$ we have

$$\begin{aligned} \square F^{\mu\nu} &= \square (\partial^\mu \Phi^\nu - \partial^\nu \Phi^\mu) = 0 \\ &= \underbrace{\partial^\mu \square \Phi^\nu}_0 - \underbrace{\partial^\nu \square \Phi^\mu}_0 = 0 \end{aligned}$$

$$\Rightarrow \left(\frac{1}{c^2} \frac{\partial^2}{\partial t^2} - \frac{\partial^2}{\partial x^2} - \frac{\partial^2}{\partial y^2} - \frac{\partial^2}{\partial z^2} \right) \underbrace{F^{\mu\nu}}_{\text{contains E and B-fields}} = 0$$

This is the wave equation, we have recovered that electromagnetic waves travel with the speed of light!