

Electromagnetic waves

In regions without any sources ("far away")
Maxwell's eqn's read

$$\vec{\nabla} \cdot \vec{E} = 0$$

$$\vec{\nabla} \times \vec{E} + \frac{\partial}{\partial t} \vec{B} = 0$$

$$\vec{\nabla} \cdot \vec{B} = 0$$

$$\vec{\nabla} \times \vec{B} - \mu_0 \epsilon_0 \frac{\partial}{\partial t} \vec{E} = 0$$

taking the curl of Faraday's law gives

$$\vec{\nabla} \times (\vec{\nabla} \times \vec{E}) + \frac{\partial}{\partial t} (\vec{\nabla} \times \vec{B}) = 0$$

$$\vec{\nabla} (\underbrace{\vec{\nabla} \cdot \vec{E}}_{=0}) - \vec{\nabla}^2 \vec{E} + \mu_0 \epsilon_0 \frac{\partial^2}{\partial t^2} \vec{E} = 0$$

$$\Rightarrow \vec{\nabla}^2 \vec{E} = \mu_0 \epsilon_0 \frac{\partial^2}{\partial t^2} \vec{E} = \frac{1}{c^2} \frac{\partial^2}{\partial t^2} \vec{E}$$

or using $\partial_\mu = (\frac{1}{c} \frac{\partial}{\partial t}, \vec{\nabla})$; $\partial^\mu \vec{E}_\mu = 0$

similarly $\vec{\nabla}^2 \vec{B} = \mu_0 \epsilon_0 \frac{\partial^2}{\partial t^2} \vec{B}$

- the wave eqn in 3 space dimensions
with plane wave solution

$$\vec{E}(\vec{r}, t) = \vec{E}_\pm e^{i(\vec{k} \cdot \vec{r} \pm \omega t)}$$

↑ wave-vector
↑ angular frequency
↑ complex amplitude

check: $\vec{\nabla}^2 \vec{E} = -k^2 \vec{E}$

$$\frac{1}{c^2} \frac{\partial^2}{\partial t^2} \vec{E} = -\frac{1}{c^2} \omega^2 \vec{E}$$

true if $|\vec{k}| = k = \frac{\omega}{c} = \frac{2\pi}{\lambda}$

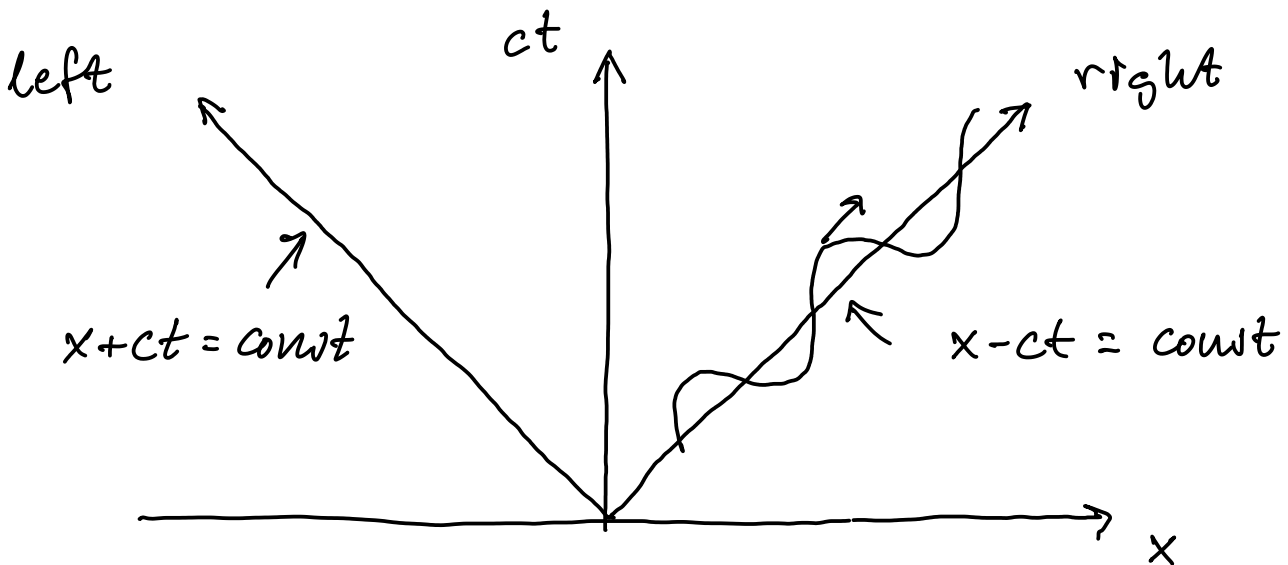
↑ wave-number ↑ propagation speed

The wave moves along so called light-cones

Choose coordinate system such that the wave moves along the x-axis ($\vec{k} = (k, 0, 0)$).

Then we have

$$kx - \omega t = k(x - ct)$$



note: the solutions are complex (simplifies the formalism) - the physical fields are given by the real parts, explicitly

$$\begin{aligned} \text{Re}(\vec{E}) &= \text{Re}(|\vec{E}| e^{i(\vec{k} \cdot \vec{r} - \omega t + \phi)}) \hat{n} \\ &= |\vec{E}| \cos(\vec{k} \cdot \vec{r} - \omega t + \phi) \hat{n} \end{aligned}$$

↑ phase

notes: in the general case a wave has several frequencies - a wave-packet

$$\vec{E} = \int \vec{E}_{\pm}(\vec{k}) e^{i(\vec{k} \cdot \vec{r} \pm \omega t)} d^3 \vec{k}$$

↑ fourier components

but thanks to the linearity of Maxwell's eqs it is often enough to consider one frequency at a time

note: a plane wave extends to infinity in the transverse directions

Writing the Laplace operator in spherical coordinates (r, θ, φ) and assuming spherical symmetry we have

$$\begin{aligned}\nabla^2 f &= \frac{1}{r^2} \frac{\partial}{\partial r} \left(r^2 \frac{\partial}{\partial r} f \right) = \frac{1}{r^2} 2r \frac{\partial}{\partial r} f + \frac{\partial^2}{\partial r^2} f \\ &= \frac{1}{r} \frac{\partial^2}{\partial r^2} (r f)\end{aligned}$$

which gives

$$\frac{1}{r} \frac{\partial^2}{\partial r^2} (r \vec{E}) = \frac{1}{c^2} \frac{\partial^2}{\partial t^2} \vec{E}$$

$$\Rightarrow \frac{\partial^2}{\partial r^2} (r \vec{E}) = \frac{1}{c^2} \frac{\partial^2}{\partial t^2} (r \vec{E})$$

with solutions

$$r \vec{E}(\vec{r}, t) = \vec{E}_{\pm} e^{i(kr \pm \omega t)}$$

or

$$\vec{E}(\vec{r}, t) = \vec{E}_{\pm} \frac{e^{i(kr \pm \omega t)}}{r}$$

spherical waves moving $\left\{ \begin{array}{l} \text{into } (+) \\ \text{out from } (-) \end{array} \right\}$ the origin

will come back to this when we consider dipole radiation (and scattering)

A wave propagates in the $\hat{k} = \frac{\vec{k}}{k}$ direction

Assuming that

$$\vec{E} = \vec{E} e^{i(\vec{k} \cdot \vec{r} - \omega t)} \hat{n}$$

Gauss' law $\vec{\nabla} \cdot \vec{E} = 0$ gives $\hat{k} \cdot \hat{n} = 0$

Faraday's law $\vec{\nabla} \times \vec{E} = -\frac{\partial}{\partial t} \vec{B}$ gives

$$i \vec{k} \times \vec{E} e^{i(\vec{k} \cdot \vec{r} - \omega t)} \hat{n} = i \omega \vec{B} e^{i(\vec{k} \cdot \vec{r} - \omega t)}$$

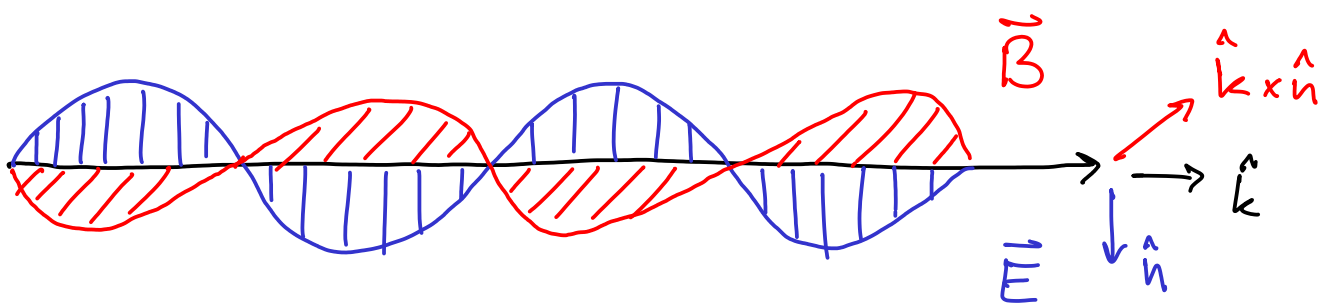
$$\Rightarrow \vec{B} = \frac{k}{\omega} \vec{E} \hat{k} \times \hat{n} e^{i(\vec{k} \cdot \vec{r} - \omega t)}$$

\uparrow
 $\frac{1}{c}$

note: $\vec{B} \cdot \hat{k} = 0$, $\vec{B} \cdot \vec{E} = 0$

$$|\vec{B}| = \frac{1}{c} |\vec{E}|$$

note: \vec{B} - field contains no extra information (same phase, speed)



called linearly polarised wave

$(\hat{k}, \hat{n}, \hat{k} \times \hat{n})$ right-handed system

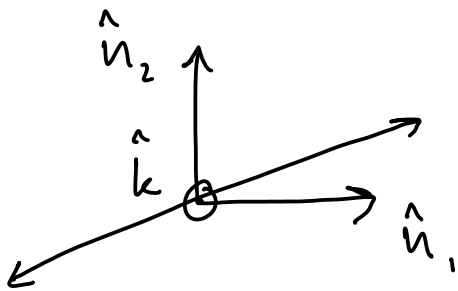
In the general case we can write \vec{E} as

$$\vec{E} = \hat{n}_1 |\tilde{E}_1| e^{i(\vec{k} \cdot \vec{r} - \omega t + \delta_1)} + \hat{n}_2 |\tilde{E}_2| e^{i(\vec{k} \cdot \vec{r} - \omega t + \delta_2)}$$

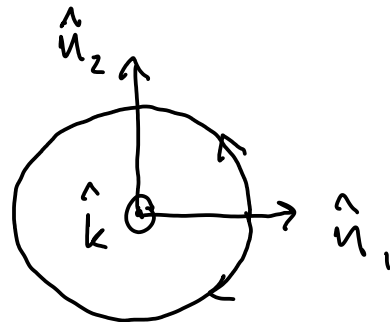
where $\hat{k} \cdot \hat{n}_1 = \hat{k} \cdot \hat{n}_2 = 0$ and preferentially $\hat{n}_1 \cdot \hat{n}_2 = 0$

If $\delta_1 = \delta_2$ then we have linear polarisation

" $\delta_1 = \delta_2 \pm \frac{\pi}{2}$ and $|\tilde{E}_1| = |\tilde{E}_2|$, circular — —



linear



circular

Energy transported by plane wave:

energy densities:

$$u_E = \frac{1}{2} \epsilon_0 [\text{Re}(\tilde{\vec{E}})]^2, \quad u_M = \frac{1}{2} \frac{1}{\mu_0} [\text{Re}(\tilde{\vec{B}})]^2$$

$$|\tilde{\vec{B}}| = \frac{1}{c} |\tilde{\vec{E}}|, \quad \frac{1}{c^2} = \mu_0 \epsilon_0$$

$$\Rightarrow u_M = u_E = \frac{1}{2} \epsilon_0 |\tilde{\vec{E}}|^2 \cos^2(\vec{k} \cdot \vec{r} - \omega t + \phi)$$

$$u = u_E + u_M = 2u_E$$

Poynting's vector: $[\hat{n} \times (\hat{k} \times \hat{n}) = \hat{k}]$

$$\vec{S} = \frac{1}{\mu_0} (\vec{E} \times \vec{B}) = \underbrace{c \epsilon_0 |\tilde{\vec{E}}|^2 \cos^2(\vec{k} \cdot \vec{r} - \omega t + \phi)}_u \hat{k}$$

We also get the time averages, $\langle \cos^2(\cdot) \rangle = \frac{1}{2}$

$$\langle u \rangle = \frac{1}{2} \epsilon_0 |\tilde{\vec{E}}|^2$$

$$\langle \vec{S} \rangle = \frac{1}{2} c \epsilon_0 |\tilde{\vec{E}}|^2 \hat{z} = c \langle u \rangle \hat{z}$$

and the intensity, I , the average power per unit area

$$I = |\langle \vec{S} \rangle| = c \langle u \rangle = \frac{1}{2} c \epsilon_0 |\tilde{\vec{E}}|^2$$

note: the time-average can be written as

$$\langle u_E \rangle = \left\langle \frac{1}{2} \epsilon_0 \vec{E}^2 \right\rangle =$$

$$\left[\vec{E}^2 = \left(\text{Re}(\tilde{\vec{E}} e^{i(\vec{k} \cdot \vec{r} - \omega t)}) \right)^2 = \frac{1}{4} \left(\tilde{\vec{E}} e^{i(\vec{k} \cdot \vec{r} - \omega t)} + \tilde{\vec{E}}^* e^{-i(\vec{k} \cdot \vec{r} - \omega t)} \right)^2 \right]$$

$$= \frac{1}{2} \epsilon_0 \frac{1}{4} \left\langle \tilde{\vec{E}}^2 e^{-2i\omega t} + \tilde{\vec{E}}^{*2} e^{2i\omega t} + 2\tilde{\vec{E}}\tilde{\vec{E}}^* \right\rangle$$

$$= \frac{1}{4} \epsilon_0 \tilde{\vec{E}}\tilde{\vec{E}}^* = \frac{1}{4} \epsilon_0 |\tilde{\vec{E}}|^2$$

Wave-propagation in linear media:

Maxwell's eqn's in matter with no free charge and no free currents

$$\begin{aligned}\vec{\nabla} \cdot \vec{D} &= 0 & \vec{\nabla} \times \vec{E} &= -\frac{\partial}{\partial t} \vec{B} \\ \vec{\nabla} \cdot \vec{B} &= 0 & \vec{\nabla} \times \vec{H} &= \frac{\partial}{\partial t} \vec{D}\end{aligned}$$

for linear (homogeneous, isotropic) media

$$\begin{aligned}\vec{D} &= \epsilon \vec{E} & \vec{B} &= \mu \vec{H} \\ &\uparrow \epsilon_0 \epsilon_r & &\uparrow \mu_0 \mu_r\end{aligned}$$

we get same eqns as in vacuum:

$$\begin{aligned}\vec{\nabla} \cdot \vec{E} &= 0 & \vec{\nabla} \times \vec{E} &= -\frac{\partial}{\partial t} \vec{B} \\ \vec{\nabla} \cdot \vec{B} &= 0 & \vec{\nabla} \times \vec{B} &= \underbrace{\mu \epsilon}_{\mu_r \epsilon_r \mu_0 \epsilon_0} \frac{\partial}{\partial t} \vec{E}\end{aligned}$$

only difference - speed of propagation

$$v = \frac{1}{\sqrt{\mu \epsilon}} = \frac{1}{\sqrt{\mu_r \epsilon_r}} \frac{1}{\sqrt{\mu_0 \epsilon_0}} = \frac{c}{n}$$

\uparrow index of refraction

as before the energy densities are

$$u_E = \frac{1}{2} \epsilon \vec{E}^2 \quad u_H = \frac{1}{2} \frac{1}{\mu} \vec{B}^2 = \frac{1}{2} \mu \vec{H}^2$$

and Poynting vector and the intensity

$$\vec{S} = \frac{1}{\mu} (\vec{E} \times \vec{B}) = \vec{E} \times \vec{H}, \quad |\vec{B}| = \frac{1}{v} |\vec{E}|$$

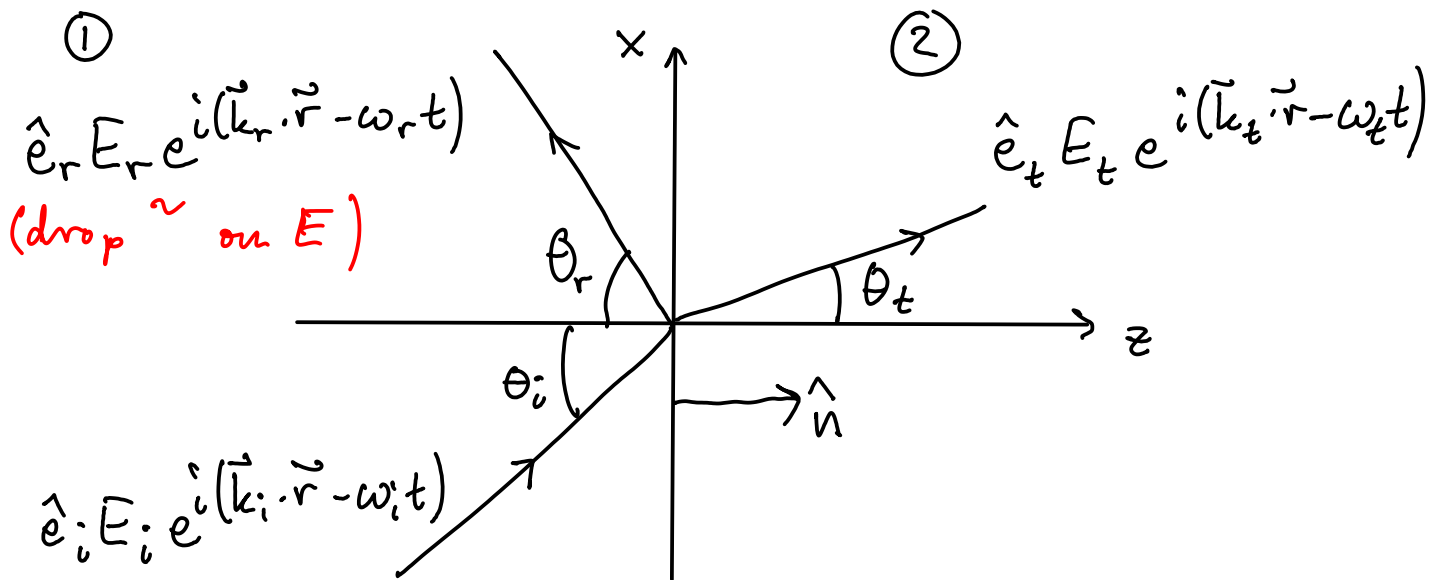
$$I = \langle |\vec{S}| \rangle = \frac{1}{2} \frac{1}{\mu v} |\vec{E}|^2 = \frac{1}{2} \epsilon v |\vec{E}|^2 = \frac{1}{2} \sqrt{\frac{\epsilon}{\mu}} |\vec{E}|^2 = \frac{1}{2} \frac{1}{Z} |\vec{E}|^2$$

$$\left[Z_0 = \sqrt{\frac{\mu_0}{\epsilon_0}} = \mu_0 c = \frac{1}{\epsilon_0 c} \approx 377 \text{ ohms}, \quad P = uI = \frac{u^2}{R} \right]$$

Reflection and transmission:

Consider light going from an optically thinner (1) medium to an optically thicker (2), i.e. $n_1 < n_2$

Plane waves for incoming, reflected and transmitted waves:



The electric fields on the two sides are

$$\vec{E}_1 = \hat{e}_r E_r e^{i(\vec{k}_r \cdot \vec{r} - \omega_r t)} + \hat{e}_i E_i e^{i(\vec{k}_i \cdot \vec{r} - \omega_i t)}$$

$$\vec{E}_2 = \hat{e}_t E_t e^{i(\vec{k}_t \cdot \vec{r} - \omega_t t)}$$

Boundary conditions:

$$D_1^\perp - D_2^\perp = \sigma_f = 0 \quad \Rightarrow \quad \epsilon_1 E_1^\perp = \epsilon_2 E_2^\perp$$

$$B_1^\perp - B_2^\perp = 0 \quad \Rightarrow \quad B_1^\perp = B_2^\perp$$

$$\vec{E}_1^\parallel - \vec{E}_2^\parallel = 0 \quad \Rightarrow \quad \vec{E}_1^\parallel = \vec{E}_2^\parallel$$

$$\vec{H}_1^\parallel - \vec{H}_2^\parallel = \vec{K}_f \times \hat{n} = 0 \quad \Rightarrow \quad \frac{1}{\mu_1} \vec{B}_1^\parallel = \frac{1}{\mu_2} \vec{B}_2^\parallel$$

(same for real and imag. parts)

these conditions apply for all t, x, y

$$\Rightarrow \begin{cases} \omega_i = \omega_r = \omega_t = \omega \\ (\vec{k}_i \cdot \vec{r})_{z=0} = (\vec{k}_r \cdot \vec{r})_{z=0} = (\vec{k}_t \cdot \vec{r})_{z=0} \end{cases}$$

In other words all \vec{k} are in the same plane
- choose the xz -plane

$$\Rightarrow k_i \sin \theta_i = k_r \sin \theta_r = k_t \sin \theta_t \quad (\text{x-components are the same})$$

$$\text{since } v_i = v_r = v_1 = \frac{c}{n_1}, \quad v_t = v_2 = \frac{c}{n_2} \quad \text{and } k v = \omega$$

we also get

$$\left[\begin{aligned} v_i = v_r &\Rightarrow k_i = k_r \Rightarrow \theta_i = \theta_r \\ \frac{\omega}{v_i} \sin \theta_i &= \frac{\omega}{v_t} \sin \theta_t \Rightarrow \frac{\omega n_1}{c} \sin \theta_i = \frac{\omega n_2}{c} \sin \theta_t \end{aligned} \right]$$

$$\begin{cases} \theta_i = \theta_r & \text{law of reflection} \\ n_1 \sin \theta_i = n_2 \sin \theta_t & \text{law of refraction} \end{cases}$$

note: these do not depend on the dynamics!

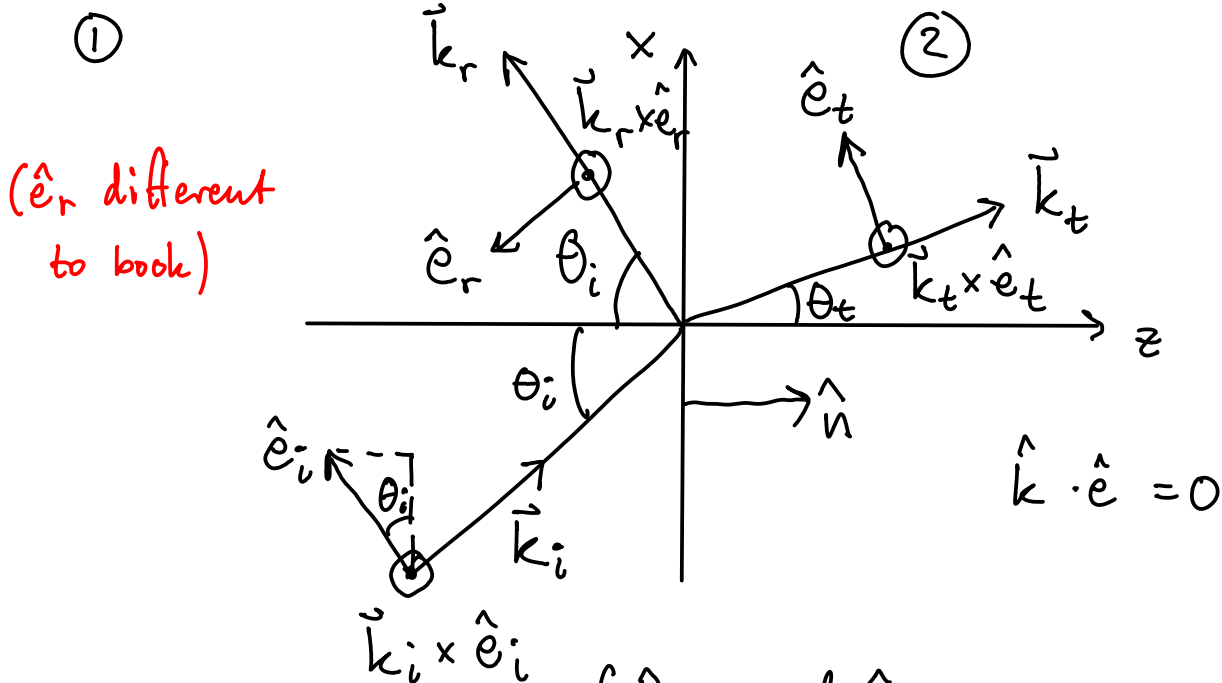
The magnetic fields:

$$\vec{B}_1 = \frac{1}{v_1} \left(E_r e^{i(\vec{k}_r \cdot \vec{r} - \omega t)} \hat{k}_r \times \hat{e}_r + E_i e^{i(\vec{k}_i \cdot \vec{r} - \omega t)} \hat{k}_i \times \hat{e}_i \right)$$

$$\vec{B}_2 = \frac{1}{v_2} E_t e^{i(\vec{k}_t \cdot \vec{r} - \omega t)} \hat{k}_t \times \hat{e}_t$$

To proceed we have to specify the polarisation:

Consider first the case when \hat{e}_i is in the xz -plane
 - called transverse magnetic (TM) polarisation



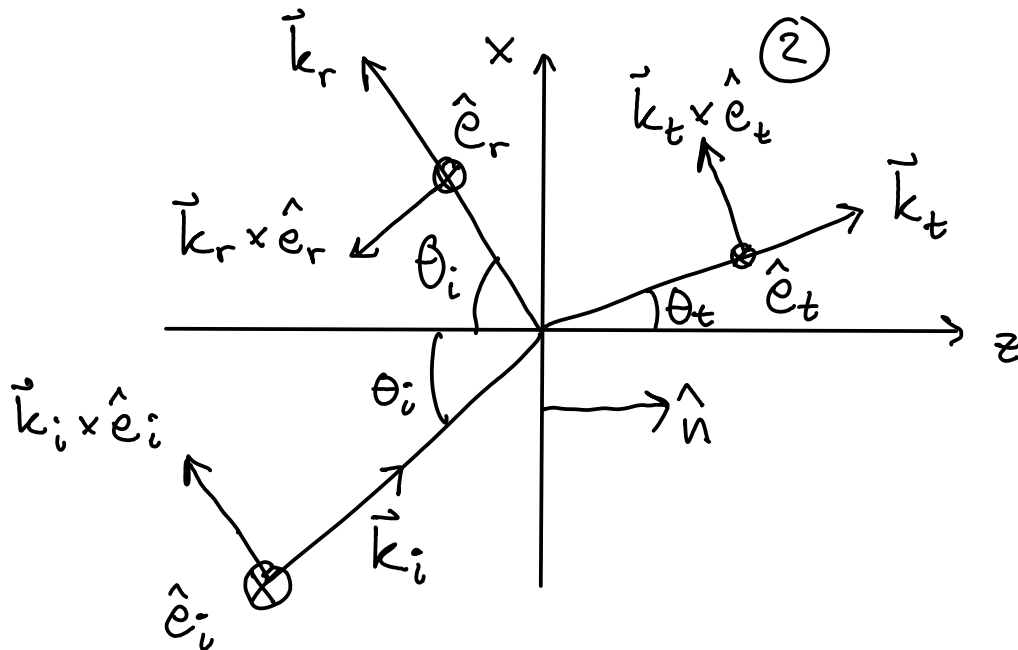
boundary cond. \Rightarrow $\begin{cases} \hat{e}_r \text{ and } \hat{e}_t \text{ in } xz\text{-plane} \\ \vec{E}^{\parallel} \text{ and } \vec{H}^{\parallel} \text{ continuous} \end{cases}$

$$\begin{cases} E_i \cos \theta_i - E_r \cos \theta_r = E_t \cos \theta_t \\ \frac{1}{\mu_1 v_1} E_i + \frac{1}{\mu_1 v_1} E_r = \frac{1}{\mu_2 v_2} E_t \end{cases} \quad \begin{aligned} |\vec{B}| &= \frac{1}{v} |\vec{E}| \\ |\vec{H}| &= \frac{1}{\mu} |\vec{B}| \end{aligned}$$

using $v = \frac{c}{n}$ the solutions are

$$\begin{cases} \frac{E_t}{E_i} = \frac{2\mu_2 n_1 \cos \theta_i}{\mu_1 n_2 \cos \theta_i + \mu_2 n_1 \cos \theta_t} \\ \frac{E_r}{E_i} = \frac{\mu_1 n_2 \cos \theta_i - \mu_2 n_1 \cos \theta_t}{\mu_1 n_2 \cos \theta_i + \mu_2 n_1 \cos \theta_t} \end{cases} \quad \text{Frenel's eqn's}$$

Consider next the case when $\hat{e}_i = -\hat{y}$
 - called transverse electric (TE) polarization
 (problem 9.17)



again \vec{E}^n and \vec{H}^n are continuous giving

$$\frac{E_t}{E_i} = \frac{2\mu_2 n_1 \cos \theta_i}{\mu_2 n_1 \cos \theta_i + \mu_1 n_2 \cos \theta_t}$$

$$\frac{E_r}{E_i} = \frac{\mu_2 n_1 \cos \theta_i - \mu_1 n_2 \cos \theta_t}{+}$$

In both cases normal incidence $\theta_i = 0$ gives

$$\frac{E_t}{E_i} = \frac{2\mu_2 n_1}{\mu_2 n_1 + \mu_1 n_2}$$

$$\frac{E_r}{E_i} = \frac{\mu_1 n_2 - \mu_2 n_1}{+} \quad (\text{up to a sign depending on how } \hat{e}_r \text{ is defined})$$

if instead $\cos \theta_i \approx 0$ we get

$$E_t \approx 0$$

$$\frac{E_r}{E_i} \approx -1$$

phase shifted with π

For TM waves the reflected wave vanishes if

$$\mu_1 n_2 \cos \theta_i = \mu_2 n_1 \cos \theta_t = \mu_2 n_1 \sqrt{1 - \frac{n_1^2}{n_2^2} \sin^2 \theta_i}$$

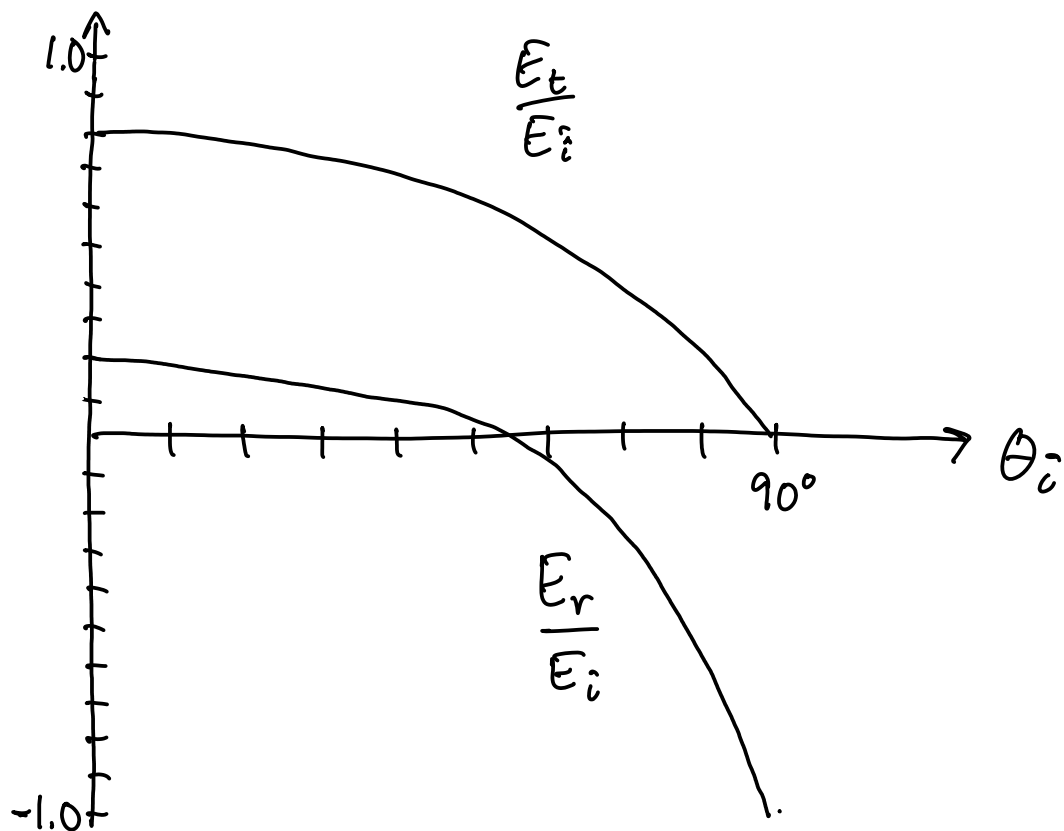
$$\left[n_1 \sin \theta_i = n_2 \sin \theta_t \right]$$

with the solution (assuming $\mu_1 = \mu_2$)

$$\tan \theta_B = \frac{n_2}{n_1} \quad \text{Brewster angle}$$

(the idea behind Polaroid shades)

Example: $\mu_1 = \mu_2 = 1$, $n_2 = 1.5$, $n_1 = 1$ gives
for TM polarisation



"Energy conservation"

To look at the energy-transport we need Poynting's vector for the two sides

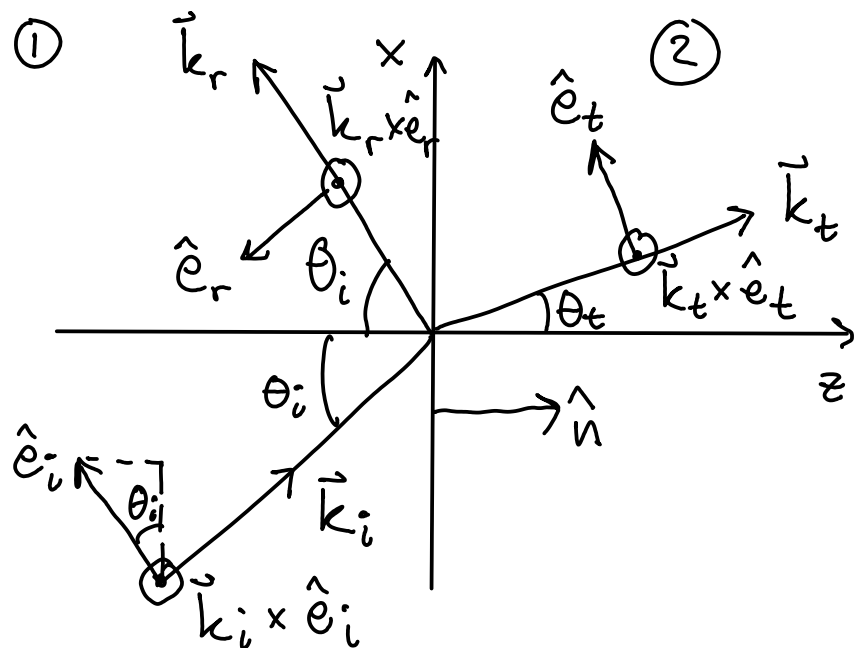
$$\vec{S}_{1,2} = \vec{E}_{1,2} \times \vec{H}_{1,2} \quad , \quad \langle \vec{S}_{1,2} \rangle = \frac{1}{2} \text{Re}(\vec{E}_{1,2} \times \vec{H}_{1,2}^*)$$

$$\vec{E}_1 = \vec{E}_i + \vec{E}_r \quad \textcircled{1}$$

$$\vec{H}_1 = \vec{H}_i + \vec{H}_r$$

$$\vec{E}_2 = \vec{E}_t$$

$$\vec{H}_2 = \vec{H}_t$$



Consider TM-polarisation

$$\vec{E}_i = \hat{e}_i E_i e^{i(\vec{k}_i \cdot \vec{r} - \omega t)} \quad , \quad \hat{e}_i = \hat{x} \cos \theta_i - \hat{z} \sin \theta_i$$

$$\vec{H}_i = \frac{1}{\mu_1 v_1} \hat{k}_i \times \hat{e}_i E_i e^{i(\vec{k}_i \cdot \vec{r} - \omega t)} \quad , \quad \hat{k}_i \times \hat{e}_i = \hat{y}$$

similarly for reflected wave but

$$\hat{e}_r = -\hat{x} \cos \theta_i - \hat{z} \sin \theta_i \quad , \quad \hat{k}_r \times \hat{e}_r = \hat{y}$$

$$\Rightarrow \langle \vec{S}_1 \rangle = \frac{1}{2\mu_1 v_1} \text{Re} \left[\underbrace{(\hat{e}_i E_i + \hat{e}_r E_r)}_{\hat{z} (E_i - E_r) \cos \theta_i + \hat{x} (E_i + E_r) \sin \theta_i} \times \hat{y} (E_i^* + E_r^*) \right]$$

$$\hat{z} (E_i - E_r) \cos \theta_i + \hat{x} (E_i + E_r) \sin \theta_i$$

$$\therefore \langle \vec{S}_1 \rangle \cdot \hat{z} = \frac{1}{2\mu_1 v_1} \text{Re} \left[(E_i^* + E_r^*) (E_i - E_r) \right] \cos \theta_i =$$

$$= \frac{1}{2\mu_1 v_1} |E_i|^2 \cos \theta_i - \frac{1}{2\mu_1 v_1} |E_r|^2 \cos \theta_i = (\langle \vec{S}_i \rangle + \langle \vec{S}_r \rangle) \cdot \hat{z}$$

for the transmitted wave we have

$$\langle \vec{S}_2 \rangle \cdot \hat{z} = \langle \vec{S}_t \rangle \cdot \hat{z} = \frac{1}{2\mu_2 v_2} |E_t|^2 \cos \theta_t$$

the boundary conditions for \vec{E}'' and \vec{H}''

$$\begin{cases} E_i \cos \theta_i - E_r \cos \theta_r = E_t \cos \theta_t \\ \frac{1}{\mu_1 v_1} E_i^* + \frac{1}{\mu_1 v_1} E_r^* = \frac{1}{\mu_2 v_2} E_t^* \end{cases}$$

then tells us that

$$\langle \vec{S}_1 \rangle \cdot \hat{z} = \langle \vec{S}_2 \rangle \cdot \hat{z}$$

$$\Rightarrow \underbrace{\frac{1}{2\mu_1 v_1} |E_i|^2 \cos \theta_i}_{I_i} = \underbrace{\frac{1}{2\mu_1 v_1} |E_r|^2 \cos \theta_r}_{I_r} + \underbrace{\frac{1}{2\mu_2 v_2} |E_t|^2 \cos \theta_t}_{I_t}$$

$$\therefore I_i = I_r + I_t$$

$$\frac{I_r}{I_i} + \frac{I_t}{I_i} = \underbrace{\frac{|E_r|^2}{|E_i|^2}}_R + \underbrace{\frac{\mu_1 v_1}{\mu_2 v_2} \frac{|E_t|^2}{|E_i|^2} \frac{\cos \theta_t}{\cos \theta_i}}_T = 1 \quad (\text{exercise})$$

when projecting onto the z-axis we can think of the energy of the incoming wave being split on the reflected and transmitted ones - but this is not true in the x-direction

So far $n_1 < n_2$. If instead $n_2 < n_1$, then Snell's law gives $\theta_t = 90^\circ$ when

$$\sin \theta_c = \frac{n_2}{n_1} \equiv \sin \theta_c$$

called critical angle (Problem 9.39)

For larger angles there is no transmitted wave

$$\cos \theta_t = \sqrt{1 - \frac{n_1^2}{n_2^2} \sin^2 \theta_i} = i \xi \quad \leftarrow \text{real}$$

$$\text{and } \left. \frac{E_r}{E_i} \right|_{TM} = \frac{\mu_1 n_2 \cos \theta_i - i \mu_2 n_1 \xi}{\mu_1 n_2 \cos \theta_i + i \mu_2 n_1 \xi}$$

$$\left. \frac{E_r}{E_i} \right|_{TE} = \frac{\mu_2 n_1 \cos \theta_i - i \mu_1 n_2 \xi}{\mu_2 n_1 \cos \theta_i + i \mu_1 n_2 \xi}$$

so that in both cases $\left| \frac{E_r}{E_i} \right|^2 = 1$

[If we consider the evanescent wave we have

$$e^{i \vec{k}_t \cdot \vec{r}} = \underbrace{e^{i |\vec{k}_t| z \cos \theta_t}}_{e^{-|\vec{k}_t| z \xi}} e^{i |\vec{k}_t| x \sin \theta_t} =$$

[Snell's law: $|\vec{k}_i| \sin \theta_i = |\vec{k}_t| \sin \theta_t$, $n_1 \sin \theta_i = n_2 \sin \theta_t$]

$$= \underbrace{e^{-|\vec{k}_i| \frac{n_2}{n_1} \xi z}}_{\text{exp. damping}} \underbrace{e^{i |\vec{k}_i| \sin \theta_i x}}_{\text{propagation in } x\text{-dir.}}$$

writing the first factor as $e^{-\frac{z}{\delta}}$ we have

$$\delta = \frac{n_1}{|\vec{k}_i| n_2 \xi} = \frac{1}{|\vec{k}_i| \sqrt{\sin^2 \theta_i - \frac{n_2^2}{n_1^2}}} = \frac{\lambda_i}{2\pi \sqrt{\sin^2 \theta_i - \sin^2 \theta_c}}$$

$$\xi = \left(\frac{n_1^2}{n_2^2} \sin^2 \theta_i - 1 \right)^{1/2} \quad |\vec{k}_i| = \frac{2\pi}{\lambda_i}$$

finally the speed of propagation becomes

$$\frac{\omega}{|\vec{k}_i| \sin \theta_i} = \frac{c}{n_1 \sin \theta_i} = \frac{v_1}{\sin \theta_i} < v_2$$

$\uparrow \sin \theta_i > \frac{n_2}{n_1}$

note that there is no energy transported into medium 2 :

$$\langle \vec{S}_2 \rangle \cdot \hat{n} = \dots = 0$$



$$\nabla^2 \vec{E} = \mu\sigma \frac{\partial}{\partial t} \vec{E} + \mu\epsilon \frac{\partial^2}{\partial t^2} \vec{E}$$

and the corresponding one for \vec{B} .

Plane wave solution

$$\vec{E}(\vec{r}, t) = \hat{e} \tilde{E} e^{i(\vec{k} \cdot \vec{r} - \omega t)}$$

with (insert in wave-equ to verify)

$$\tilde{k}^2 = \tilde{k}^2 = \mu\epsilon\omega^2 + i\mu\sigma\omega = \mu\left(\epsilon + \frac{i\sigma}{\omega}\right)\omega^2$$

↑
complex wave-vector

can also define complex index of refraction

$$\tilde{k}^2 \equiv \frac{\tilde{n}^2}{c^2} \omega^2 \Rightarrow \tilde{n}^2 = \mu_r \epsilon_r \left(1 + i \frac{\sigma}{\omega\epsilon}\right)$$

writing $\tilde{k} = k + i\alpha$ and $\tilde{n} = n + iv$ we also have

$$k = n \frac{\omega}{c}, \quad \alpha = v \frac{\omega}{c} \quad (\text{implicitly this means that } \tilde{k} = \vec{k} + i\alpha \hat{k})$$

some gymnastics with complex numbers gives

$$n^2 = \frac{\epsilon_r \mu_r}{2} \left(\sqrt{1 + \left(\frac{\sigma}{\epsilon\omega}\right)^2} + 1 \right)$$

$$n^2 - v^2 = \mu_r \epsilon_r$$

$$v^2 = \frac{\epsilon_r \mu_r}{2} \left(\sqrt{1 + \left(\frac{\sigma}{\epsilon\omega}\right)^2} - 1 \right)$$

$$2nv = \mu_r \epsilon_r \frac{\sigma}{\omega\epsilon}$$

$$(\tilde{k} = (k + i\alpha) \hat{z})$$

Considering only normal incidence (we have

$$\vec{E}(\vec{r}, t) = \hat{e} \tilde{E} \underbrace{e^{-\alpha z}} e^{i(kz - \omega t)}$$

exponential damping, skin depth $d = \frac{1}{\alpha}$

Faraday's law gives the corresponding \vec{B} -field

$$\vec{B}(\vec{r}, t) = \hat{k} \times \hat{e} \frac{\tilde{k}}{\omega} \tilde{E} e^{-\alpha z} e^{i(kz - \omega t)}$$

note: \vec{B} and \tilde{E} have different phase

$$\tilde{k} = |\tilde{k}| e^{i\varphi}, \quad \varphi = \arctan \frac{\alpha}{k}$$

Boundary conditions:

$$\epsilon_1 E_1^\perp - \epsilon_2 E_2^\perp = \sigma_f \quad \vec{E}_1^\parallel - \vec{E}_2^\parallel = 0$$

$$B_1^\perp - B_2^\perp = 0 \quad \frac{1}{\mu_1} \vec{B}_1^\parallel - \frac{1}{\mu_2} \vec{B}_2^\parallel = \vec{K}_f \times \hat{n}$$

normal incidence $\Rightarrow E_1^\perp = E_2^\perp = 0 \Rightarrow \sigma_f = 0$

ohmic conductor ($\vec{J}_f = \sigma \vec{E}$), \vec{E} finite $\Rightarrow \vec{K}_f = 0$

\therefore laws of refraction and reflection unchanged for normal incidence - only modifications

\tilde{n}_2 and \tilde{k}_t complex (and $\nu, \mu \rightarrow \infty$)

note: for perfect conductor ($\sigma \rightarrow \infty$) $n \rightarrow \infty$
 \Rightarrow total reflection (cf. mirror)

Alternatively we can define a complex $\tilde{\epsilon}_r$

$$\tilde{\epsilon}_r = \epsilon_r + i \frac{\sigma}{\omega \epsilon_0}, \quad \nabla^2 \vec{E} = \mu \epsilon_0 \tilde{\epsilon}_r \frac{\partial^2 \vec{E}}{\partial t^2}$$

instead of having free currents

Complex $\tilde{\epsilon}_r$ also arises more generally when describing dispersive materials

Dispersion:

Consider simple (classical) model for how the electrons in a material are affected by an electric field

$$\vec{F} = m_e \frac{d^2}{dt^2} \vec{r} = \underset{\substack{\uparrow \\ \text{spring const.}}}{-k \vec{r}} - \underset{\substack{\uparrow \\ \text{damping}}}{\alpha \frac{d}{dt} \vec{r}} - e \underset{\substack{\uparrow \\ \text{electric field}}}{\vec{E}}$$

rewrite as

$$\frac{d^2}{dt^2} \vec{r} + \frac{\alpha}{m_e} \frac{d}{dt} \vec{r} + \frac{k}{m_e} \vec{r} = - \frac{e}{m_e} \vec{E}$$

\Rightarrow oscillatory motion with

$$\text{eigenfrequency } \omega_0^2 = \frac{k}{m_e}$$

$$\text{damping const } \gamma = \frac{\alpha}{m_e}$$

note: magnetic force neglected

assume harmonic time dep. $\vec{E} = \vec{E}(\vec{r}) e^{-i\omega t}$

$$\Rightarrow \vec{r} = - \frac{e}{m_e} \frac{1}{\omega_0^2 - \omega^2 - i\omega\gamma} \vec{E}$$

departure from equilibrium $\langle \vec{r} \rangle = 0$ gives rise to electric dipole moment

$$\vec{p} = -e \vec{r} = \frac{e^2}{m_e} \frac{1}{\omega_0^2 - \omega^2 - i\omega\gamma} \vec{E}$$

with N molecules per unit volume and f_j electrons with eigen frequency ω_j and damping γ_j per molecule we get

the polarisation

$$\vec{p} = \frac{Ne^2}{m_e} \sum_j \frac{f_j}{\omega_j^2 - \omega^2 - i\omega\gamma_j} \quad (11)$$
$$\tilde{\chi}_e \epsilon_0$$

and the relative permittivity $\tilde{\epsilon}_r = 1 + \tilde{\chi}_e$

Defining the plasma frequency

$$\omega_p^2 = \frac{Ne^2}{\epsilon_0 m_e} \sum_j f_j, \quad \sum_j f_j = Z$$

we finally get

$$\tilde{\epsilon}_r = 1 + \frac{\omega_p^2}{Z} \sum_j \frac{f_j}{\omega_j^2 - \omega^2 - i\omega\gamma_j}$$

as before we have a complex index of refraction and setting $\mu_r = 1$ we have

$$\tilde{n}^2 = \tilde{\epsilon}_r$$

with solutions ($\tilde{n} = n + iv$)

$$n = \frac{1}{\sqrt{2}} \sqrt{\text{Re}(\tilde{\epsilon}_r) + |\tilde{\epsilon}_r|}$$

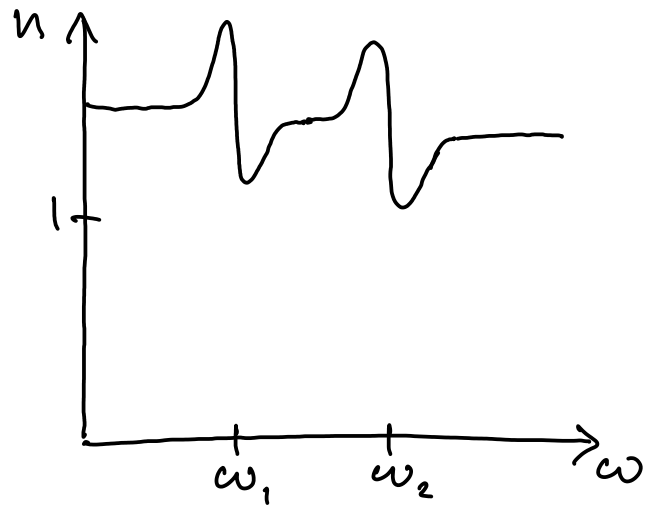
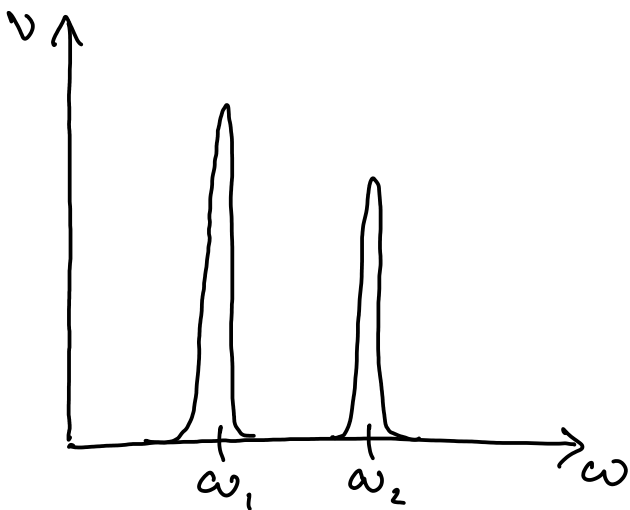
$$v = \frac{1}{2n} \text{Im}(\tilde{\epsilon}_r)$$

Now

$$\text{Re}(\tilde{\epsilon}_r) = 1 + \frac{\omega_p^2}{Z} \sum_j \frac{f_j (\omega_j^2 - \omega^2)}{(\omega_j^2 - \omega^2)^2 + \omega^2 \gamma_j^2}$$

$$\text{Im}(\tilde{\epsilon}_r) = \frac{\omega_p^2}{Z} \sum_j \frac{f_j \omega \gamma_j}{(\omega_j^2 - \omega^2)^2 + \omega^2 \gamma_j^2}$$

So $\text{Im}(\tilde{\epsilon}_r) \ll \text{Re}(\tilde{\epsilon}_r) \Rightarrow v \approx c$ unless $\omega \approx \omega_j$
 i.e. at one of the eigen frequencies



(absorption lines in spectra from stars)

note that for $\omega \gg \omega_j$ we have

$$\text{Re}(\tilde{\epsilon}_r) \approx 1 + \frac{\omega_p^2}{\omega^2} \sum_j f_j \frac{-\omega^2}{\omega^4} = 1 - \frac{\omega_p^2}{\omega^2}$$

$$\text{Im}(\tilde{\epsilon}_r) \approx \frac{\omega_p^2}{\omega^2} \sum_j f_j \frac{\omega \gamma_j}{\omega^4} \sim \frac{\omega_p^2}{\omega^3} \rightarrow 0$$

in other words $\epsilon_r < 1$ and $n = \sqrt{\epsilon_r} < 1$

$$\Rightarrow v = \frac{c}{n} > c \quad !$$

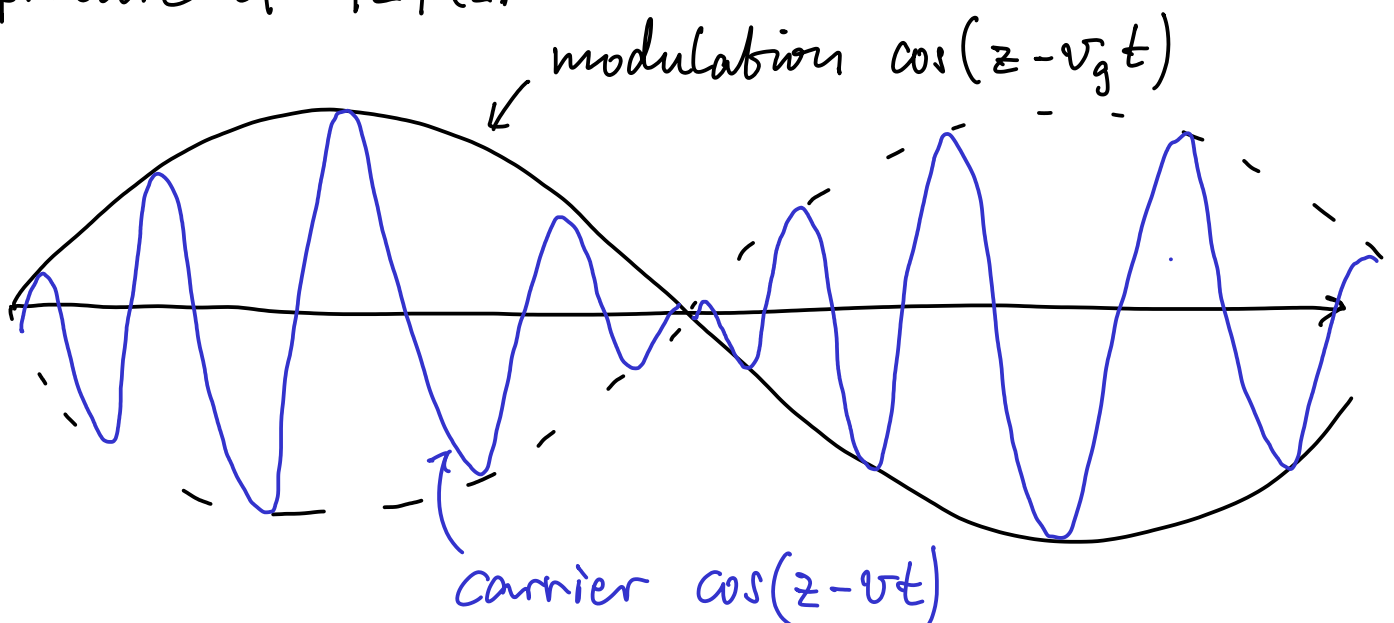
information traveling faster than light?
 No, plane wave cannot transmit info.

Consider signal consisting of two plane waves with different wave numbers and frequencies

$$\begin{aligned} \vec{E} &= \hat{e} \vec{E} \left[e^{i(k+\Delta k)z - (\omega+\Delta\omega)t} + e^{i(k-\Delta k)z - (\omega-\Delta\omega)t} \right] \\ &= \hat{e} \vec{E} \underbrace{\left[e^{i\Delta k z - \Delta\omega t} + e^{-i\Delta k z + \Delta\omega t} \right]}_{2 \cos(\Delta k z - \Delta\omega t)} e^{i(kz - \omega t)} \end{aligned}$$

signal with speed of prop. $v_g = \frac{\Delta\omega}{\Delta k}$

picture of $|\vec{E}|(z)$



More generally the group velocity is given by

$$v_g = \frac{d\omega}{dk}$$

in our case above we had

$$\epsilon_r = n^2 = 1 - \frac{\omega_p^2}{\omega^2} \Rightarrow \omega = \sqrt{\omega_p^2 + c^2 k^2}$$

$$\left[n^2 = c^2 \frac{k^2}{\omega^2} \right]$$

$$\text{so } v_g = \frac{c^2 k}{\sqrt{\omega_p^2 + c^2 k^2}} = c^2 \frac{k}{\omega} = nc < c$$

even more generally $k^2 = \frac{\omega^2}{v^2} = \frac{\omega^2}{c^2} \epsilon_r$ gives

$$v_g = \left[\frac{dk}{d\omega} \right]^{-1} = \left[\frac{1}{v} \frac{k}{\omega} \frac{d\omega}{d\omega} \right]^{-1} =$$

$$= v \left[\frac{dk^2}{d\omega^2} \right]^{-1} = v \left[\frac{\epsilon_r}{c^2} + \frac{\omega^2}{c^2} \frac{d\epsilon_r}{d\omega} \right]^{-1}$$

$$= \frac{c}{\sqrt{\epsilon_r} + \frac{\omega^2}{\sqrt{\epsilon_r}} \frac{d\epsilon_r}{d\omega^2}} = \frac{c}{\sqrt{\epsilon_r} + \omega \frac{d\sqrt{\epsilon_r}}{d\omega}}$$